

 \cdot HFO-brown-carbon λ -dependent imaginary refractive indices and mass absorption cross-sections are much higher than previously estimated. cross-sections are much higher than previously estimated.

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Abstract

 We characterized the chemical composition and optical properties of particulate mat- ter (PM) emitted by a marine diesel engine operated on heavy fuel oil (HFO), marine gas oil (MGO) and diesel fuel (DF). For all three fuels, ∼ 80% of submicron PM was organic (and sulfate, for HFO at higher engine loads). Emission factors varied only slightly with ³⁷ engine load. Refractory black carbon (rBC) particles were not thickly coated for any fuel; rBC was therefore externally mixed from organic and sulfate PM. For MGO and DF PM, ³⁹ rBC particles were lognormally distributed in size (mode at $d_{rBC} \approx 120$ nm). For HFO, much larger rBC particles were present. Combining the rBC mass concentrations with in-⁴¹ situ absorption measurements yielded an rBC mass absorption coefficient $MAC_{BC,780 \text{ nm}}$ $\frac{42}{4}$ of 7.8 ± 1.8 m² g⁻¹ at 780 nm for all three fuels. Using positive deviations of the absorp-⁴³ tion Ångström exponent (AAE) from unity to define brown carbon (brC), we found that 44 brC absorption was negligible for MGO or DF PM (AAE(370,880 nm) $\approx 1.0 \pm 0.1$) but 45 typically 50 % of total 370 nm absorption for HFO PM. Even at 590 nm, $\sim 20\%$ of the total absorption was due to brC. Using absorption at 880 nm as a reference for BC absorption and normalizing to organic PM mass, we obtained a MAC_{OM,370 nm} of $0.4 \text{ m}^2 \text{ g}^{-1}$ at typical operating conditions. Furthermore, we calculated an imaginary refractive index of $(0.045 \pm 0.025)(\lambda/370 \text{ nm})^{-3}$ for HFO PM at 370 nm> $\lambda > 660 \text{ nm}$, more than twofold greater than previous recommendations. Climate models should account for this substan-tial brC absorption in HFO PM.

Plain language summary (JGR feature)

 We characterized the fundamental properties of marine-engine exhaust that are rel- evant to its aerosol-radiation interactions in climate models. In particular, we focussed on "brown carbon" light absorption (i.e., absorption in excess of that expected for the black ₅₆ carbon in canonical soot). We found that brown carbon can increase the 370 nm direct ra-diative forcing of heavy-fuel-oil marine exhaust by 18 % over snow.

1 Introduction

 Ship engines are the major source of combustion aerosols in the marine environ- ment, where almost no other anthropogenic pollution sources exist. Understanding the ⁶¹ climate-relevant optical properties of this particulate matter (PM) is therefore essential to ⁶² the accurate determination of anthropogenic radiative effects on the marine atmosphere [\[](#page-32-2)*[Endresen](#page-30-0)*, [2003;](#page-30-0) *[Lauer et al.](#page-32-0)*, [2007;](#page-32-0) *[Unger et al.](#page-35-0)*, [2010;](#page-35-0) *[Lindstad et al.](#page-32-1)*, [2015;](#page-32-1) *[Marelle](#page-32-2) [et al.](#page-32-2)*, [2016\]](#page-32-2). In addition, a large fraction of ship PM is emitted near shore [*[Eyring et al.](#page-30-1)*, [2010\]](#page-30-1), which has a strong potential to negatively impact human health [*[Corbett et al.](#page-29-0)*,

[2007;](#page-29-0) *[Oeder et al.](#page-33-0)*, [2015\]](#page-33-0) as well as visibility.

Ship-engine PM contains significant amounts of organic PM (OM) as well as black

carbon (BC) [*[Lack et al.](#page-31-0)*, [2009;](#page-31-0) *[Buffaloe et al.](#page-29-1)*, [2014;](#page-29-1) *[Mueller et al.](#page-33-1)*, [2015;](#page-33-1) *[Price et al.](#page-34-0)*,

[2016\]](#page-34-0). When heavy fuel oil (HFO) is used, sulfates also contribute significantly to PM

⁷⁰ mass [*[Lack et al.](#page-31-0)*, [2009;](#page-31-0) *[Popovicheva et al.](#page-34-1)*, [2009;](#page-34-1) *[Mueller et al.](#page-33-1)*, [2015\]](#page-33-1). Typically, BC con-

 71 tributes a relatively small fraction to the mass of HFO PM. The overall single-scattering

 72 albedo (SSA) of HFO PM aerosols is therefore relatively high, such that ship-engine PM

⁷³ may result in an overall cooling over the (low-albedo) ocean surface [*[Lauer et al.](#page-32-0)*, [2007;](#page-32-0)

⁷⁴ *[Unger et al.](#page-35-0)*, [2010\]](#page-35-0), but a localized warming over high-albedo surfaces or on snow and ice

⁷⁵ [*[Lack and Corbett](#page-31-1)*, [2012\]](#page-31-1). The accurate quantification of these radiative effects requires an

⁷⁶ accurate measurement of the light-absorption properties of both BC and OM.

 π Light-absorbing OM has previously been inferred to contribute significantly to total ⁷⁸ HFO PM absorption at shorter visible wavelengths λ, based on the observation of 2-λ absorption Ångström exponents (AAE(370, 880 nm); Eq. 1) much greater than unity [*Muelle* ⁷⁹ [s](#page-33-1)orption Ångström exponents (AAE(370, 880 nm); Eq. [1\)](#page-7-0) much greater than unity [*[Mueller](#page-33-1)* ⁸⁰ *[et al.](#page-33-1)*, [2015\]](#page-33-1). Accounting for this light-absorbing OM in radiative models of the Earth's at-81 mosphere requires knowledge of intensive optical properties, such as imaginary refractive ⁸² index (k_{OM}) and mass absorption cross-section (MAC_{OM}), which have not previously been 83 reported.

 84 A major goal of the present work was to quantify k_{OM} and MAC_{OM} for use in cli-⁸⁵ mate models. This quantification was achieved by applying the popular [*[Fialho et al.](#page-30-2)*, ⁸⁶ [2005;](#page-30-2) *[Sandradewi et al.](#page-34-2)*, [2008;](#page-34-2) *[Favez et al.](#page-30-3)*, [2009;](#page-30-3) *[Grenfell et al.](#page-30-4)*, [2011;](#page-30-4) *[Bahadur et al.](#page-29-2)*, ⁸⁷ [2012;](#page-29-2) *[Cazorla et al.](#page-29-3)*, [2013;](#page-29-3) *[Lack and Langridge](#page-31-2)*, [2013;](#page-31-2) *[Yuan et al.](#page-35-1)*, [2016;](#page-35-1) *[Pokhrel et al.](#page-33-2)*, 88 [2017;](#page-33-2) *[Zotter et al.](#page-36-0)*, [2017\]](#page-36-0) two-component "BC+brC" model: (i) light absorption in the 89 near infrared ($\lambda \geq 780 \text{ nm}$) is attributed to BC only, (ii) BC absorption is assumed to ⁹⁰ depend inversely on λ (AAE_{BC} = 1), and (iii) when measured visible-light absorption ex-91 ceeds predicted absorption by BC, the excess is attributed to light-absorbing OM (so- $_{92}$ called brown carbon or brC). In this context, we define BC as refractory, light-absorbing 93 carbon with an AAE of unity between $370 \text{ nm} < \lambda < 950 \text{ nm}$ [*[Petzold et al.](#page-33-3)*, [2013\]](#page-33-3). We 94 consider BC as the sole light-absorbing species at $\lambda \ge 780$ nm and, when the aerosol AAE exceeds unity, attribute the excess absorption to brC. While brC is therefore defined purely exceeds unity, attribute the excess absorption to brC. While brC is therefore defined purely ⁹⁶ from the optical properties of the aerosol, we also calculate intensive brC properties by 97 normalizing brC absorption to total organic PM mass.

 For HFO PM, the two-component BC+brC absorption approach may not fully de- scribe the aerosol. First, HFO exhaust may contain asphaltenes, large polyaromatic molecules that may contribute to PM light absorption in the near infrared, besides BC [*[Mullins](#page-33-4)*, [2010\]](#page-33-4). Second, light-absorbing char particles may also be present in HFO PM [*[Lyyrä-](#page-32-3) [nen et al.](#page-32-3)*, [1999;](#page-32-3) *[Chen et al.](#page-29-4)*, [2005;](#page-29-4) *[Popovicheva et al.](#page-34-1)*, [2009\]](#page-34-1). These char particles are formed when fuel droplets graphitize rather than vapourize at the high flame temperatures [*[Linak et al.](#page-32-4)*, [2000\]](#page-32-4) and are therefore a subcategory of BC, with typical physical diameters of ∼ 1 μm. Their large physical diameter places char particles outside of the Rayleigh regime, potentially invalidating the AAE_{BC} = 1 assumption of the BC+brC approach. Unregime, potentially invalidating the $AAE_{BC} = 1$ assumption of the BC+brC approach. Un- fortunately, our present measurements do not provide sufficient information to address these char particles explicitly. To the extent that $AAE_{char} \neq 1$, our approach may over- estimate or underestimate the absorption properties of brC. However, even if this was the case, our two-species (BC+brC) approach would still provide the optical parameters re quired for climate models to correctly represent the wavelength-dependent light absorption of HFO PM which we have observed; the inaccuracy would be in the relative attribution [o](#page-34-3)f absorption to the different light-absorbing species. Moreover, as demonstrated by *[Saleh](#page-34-3) [et al.](#page-34-3)* [\[2016\]](#page-34-3), climate models which employ Mie theory in radiative transfer calculations produce more accurate estimates of radiative forcing when Mie-theory-retrieved parame-ters are used, compared to more-complex retrievals.

 In the following, we determine the optical properties of brC from a ship engine operating on three different fuels: HFO, marine gas oil (MGO), and EN 590 diesel fuel (DF). The upcoming sections discuss (1) speciated PM emission factors and the PM mix- ing state, (2) the PM SSA and MAC_{BC} at 780 nm obtained by the extinction-minus-scattering technique combined with refractory BC (rBC) mass concentrations measured by a sin- gle particle soot photometer (SP2) and (3) brC absorption using the BC+brC conceptual 123 model in terms of MAC_{OM} and k_{OM} by using OM concentrations measured by an aerosol mass spectrometer (AMS). Note that the analysis in part (3) does not incorporate the SP2 measurements.

2 Methods

2.1 Experimental

 The engine used in this study was a single-cylinder research engine installed at the Institute of Piston Machines and Internal Combustion Engines at the University of Ros- tock in Germany. The size of the combustion chamber and layout of the engine is typi-¹³¹ cal of engines used on smaller ships as a main power supply, on large ships for ancillary power, or as a backup power supply on land, e.g. in hospitals. The engine is of a four- stroke, single-cylinder design, with a 150 mm bore and 180 mm stroke and operates at a nominal 1500 rpm with a maximum power of 80 kW. Previous publications have provided further details on the engine [*[Oeder et al.](#page-33-0)*, [2015;](#page-33-0) *[Streibel et al.](#page-35-2)*, [2017\]](#page-35-2) and reported de- tailed characterizations of its emissions [*[Mueller et al.](#page-33-1)*, [2015\]](#page-33-1), however, we emphasize that ¹³⁷ the engine was reconfigured between these two studies. During this study, certain engine operating parameters were varied to investigate their influence on the emissions; none of these parameters were identified as having influenced the properties discussed herein. The measurements took place between November $10th$ and December $9th$, 2014.

 The sampling system was similar to that used by *[Streibel et al.](#page-35-2)* [\[2017\]](#page-35-2) and is shown in Fig. [1.](#page-23-0) Emissions (sampled 1.5 m from the engine at > 500 °C) were first passed through $\frac{1}{2}$ consistent that $300 °C$ heated lines to a pre-cyclone with a nominal cut-off aerodynamic diameter of 2.5 μ m, then diluted by a factor of roughly 12 in a two-stage dilutor. The two-stage dilutor com- bined a porous-tube stage, where compressed air flowed through pores in a cylinder to provide a sheath flow and minimize wall losses, with an ejector dilutor, at a flow rate of 150 L min^{-1} . The exact dilution ratio was measured by online CO₂ monitors. After this dilution stage, filter samples for thermal/optical reflectance analysis (resulting in NIOSH 5040 EC/OC data) and elemental analysis (discussed elsewhere; *[Corbin et al.](#page-29-5)* [2017\)](#page-29-5) were

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 taken. Subsequently, two tenfold ejector dilutors (Dekati Ltd., Finland) were employed to reduce PM concentrations to atmospheric levels and ambient temperatures. The steel ¹⁵² sampling line (10 mm i.d.) to the aethalometer was approximately 5 m long with a total flow of 4.5 L min⁻¹, while the other online instruments sampled at a total of 13 L min⁻¹ through first ~7 m section of tubing and then at $3 L \text{ min}^{-1}$ through a second 3 m section of tubing. This flow was then divided between similar lengths of line for the instruments.

 The majority of the instruments sampled at high time resolution (1 min or faster), ¹⁵⁷ with the exception of the particle-filter samples. For the data presented below, the 30 minute filter samples were generally used to define periods of interest (after manual inspection) and data were averaged over these periods.

 Particle size distributions were obtained by a Scanning Mobility Particle Sizer (SMPS; ¹⁶¹ ⁸⁵Kr bipolar charger; PSI-constructed DMA equivalent to the model 3081 of TSI Inc., USA; TSI CPC 3022A) and Fast Mobility Particle Sizer (FMPS, model 3091, TSI Inc.). The SMPS measurements are considered more reliable as data-inversion issues (influenc- ing both sizing and counting) have been reported for the FMPS [*[Zimmerman et al.](#page-36-1)*, [2015\]](#page-36-1). We therefore used the SMPS data to obtain geometric standard deviations (GSD) of the size distribution for the Mie model. However, due to limited coverage of these SMPS data, the FMPS data were used to obtain geometric mean particle diameters (GMDs). These were retrieved from the FMPS data by fitting a bimodal lognormal function to the data using custom code written in Igor Pro (version 6.32, WaveMetrics, OR, USA). The FPMS data were first corrected using the routine published by *[Zimmerman et al.](#page-36-1)* [\[2015\]](#page-36-1), 171 which uses a separate CPC 3022A as a reference for the overall number concentration. The GMD obtained from the corrected FMPS data is considered more reliable than the FMPS GSD [*[Zimmerman et al.](#page-36-1)*, [2015\]](#page-36-1).

 For some samples, aerosol particle mass was measured for mobility-size-selected particles using a home-built DMA and a CPC (Model 3022A, TSI Inc. USA) sampling behind an Aerosol Particle Mass analyzer (APM Model 3601, Kanomax Japan). The DMA-177 APM-CPC data was analyzed as described in *[Zieger et al.](#page-35-3)* [\[2017\]](#page-35-3) to yield particle effec- tive density measurements for the ultrafine mode, which was dominated by organics and 179 sulfate. Considering that the organic–sulfate ultrafine particles were liquid and therefore spherical, these effective densities correspond directly to particle densities. Effective den-181 sity measurements of rBC were attempted, but were not feasible as the rBC number con-centrations were much lower.

2.2 Single Particle Soot Photometer (SP2)

 A Single Particle Soot Photometer (SP2; Droplet Measurement Technologies, CO, USA) was employed to measure black carbon concentrations by laser-induced incandes- cence [*[Stephens et al.](#page-35-4)*, [2003;](#page-35-4) *[Schwarz et al.](#page-34-4)*, [2006\]](#page-34-4). Data were analyzed using the PSI SP2 Toolkit, version 4.112. The SP2 brings BC-containing particles to incandescence during their passage through a continuous-wave, intracavity, 1064 nm laser. The instrument detects rBC cores with mass (or volume-equivalent size, d_{rBC} , considering a void-free mate-

rial density of 1800 kg m^3) from ~ 0.7 fg (~ 80 nm) to ~ 200 fg (~ 600 nm). For parti- cles in this size range, the total rBC mass reported by SP2 has been validated as accurate by multiple independent studies *[Slowik et al.](#page-35-5)* [\[2007\]](#page-35-5); *[Kondo et al.](#page-31-3)* [\[2011\]](#page-31-3); *[Laborde et al.](#page-31-4)* [\[2012a\]](#page-31-4). The total mass concentration reported below is corrected for the mass fraction [o](#page-33-3)utside of this range (Section [3.1.1\)](#page-11-0). We use the term "rBC" as recommended by *[Petzold](#page-33-3)* ¹⁹⁵ *[et al.](#page-33-3)* [\[2013\]](#page-33-3) whenever reporting SP2 data in a quantitative manner, e.g. C_{rBC} for BC mass 196 concentrations and d_{rBC} for mass-equivalent diameters. We use the term rBC core to clar- ify that any non-refractory, internally-mixed material is not measured by incandescence as it vaporizes well below the ∼ 4000 K sublimation point of BC.

 The SP2 also collects light-scattering signals, acting as an optical sizer for BC-free 200 particles ($d_{\text{minimum}} \approx 160 \text{ nm}$). For BC-containing particles, the scattering signal requires ₂₀₁ careful data processing because laser heating causes coatings to evaporate during measure- ment. This evaporation is observable in the time-resolved scattering signals. For uncoated or moderately-coated particles, peak scattering signals occur when particles are relatively closer to the centre of the Gaussian profile of the SP2 laser (since incident light inten- sity is highest closer to the centre) and simultaneously with particle incandescence. For thickly coated particles, this peak is observed substantially earlier and prior to particle incandescence, corresponding to the substantial reduction in particle volume when coating material evaporates. This binary mixing-state classification is commonly referred to as "delay-time analysis" [*[Moteki and Kondo](#page-33-5)*, [2007\]](#page-33-5). On this basis, particles can be classified as either 'thickly-coated' (significant difference between peak scattering and incandescence signals) or 'moderately- or uncoated' (no significant difference). The lower detection limit for this classification is higher than that for rBC quantification, since the SP2 scattering measurement is less sensitive than the incandescence measurement. The upper limit for ²¹⁴ this classification is imposed by saturation of the scattering signal detector, which occurs ²¹⁵ for large overall particle sizes. We present data only for particles within these limits. Note that a particle consisting of an rBC particle coagulated with other material may appear to be 'thickly-coated' in the SP2, resulting in an overestimated fraction of thickly-coated particles. Note also that a precise distinction between 'moderate or no coating' and 'thick coating', for example in terms of BC volume fraction, cannot be given without knowledge of the thermochemical properties of the coating [*[Sedlacek et al.](#page-34-5)*, [2015\]](#page-34-5). Based on our pre- vious comparisons of coating-thickness to delay-time data for atmospheric BC particles (unpublished work), we estimate that 'moderate or no coating' describe particles contain-ing 70–100 % rBC volume.

 A more complex data analysis approach allows the coating thickness of BC-containing particles to be retrieved quantitatively, as follows. By retrieving the scattering signal at 3 % of the maximum laser intensity, at which point no coating material has yet evap- orated, the scattering cross-section of the total, possibly-coated, particle can be deter- mined [*[Gao et al.](#page-30-5)*, [2007;](#page-30-5) *[Taylor et al.](#page-35-6)*, [2015\]](#page-35-6). Conversely, by retrieving the scattering sig- nal after coatings have evaporated but before the onset of incandescence, the scattering cross-section of uncoated rBC can be determined [*[Laborde et al.](#page-31-5)*, [2012b\]](#page-31-5). The latter value can be used to calculate the rBC-mass-equivalent volume of the rBC core, given a com-

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232 plex refractive index of the core $m_{\text{core}} = (n, k)$. The precise value of m_{core} is not well con-
233 strained and may vary between BC materials, but the empirical relationship $k \approx (n-1)$ has strained and may vary between BC materials, but the empirical relationship $k \approx (n-1)$ has ²³⁴ been shown by *[Bond and Bergstrom](#page-29-6)* [\[2006\]](#page-29-6), as discussed further by *[Moteki et al.](#page-33-6)* [\[2010\]](#page-33-6). 235 Employing this constraint, we determined m_{core} as $1.9+0.8i$ by fitting the rBC-core volume ²³⁶ determined by scattering to that determined by incandescence. Our approach of adjusting m_{core} only serves as an internal calibration of the coating-thickness retrieval, and must not ²³⁸ be misinterpreted as a measurement of the BC refractive index, which would require addi- $_{239}$ tional independent measurements. Our value of m_{core} is smaller than that typically used in ²⁴⁰ SP2 analysis $(2.26 + 1.26i)$, but is consistent with that used by *[Laborde et al.](#page-31-4)* [\[2012a\]](#page-31-4) for
²⁴¹ propane-flame soot. Using the typical m_{core} value of 2.26 + 1.26*i* corresponds to smaller propane-flame soot. Using the typical m_{core} value of $2.26 + 1.26i$ corresponds to smaller ²⁴² rather than larger coating thicknesses in the analysis presented below. We note that it is ²⁴³ typical to observe some negative coating thicknesses in this analysis due to random noise ²⁴⁴ on the single-particle level. Investigations in our laboratory have found that this noise is ²⁴⁵ due mainly to the variability of particle velocity through the laser beam and uncertainty in 246 determining the scattering signal at 3% of the maximum laser intensity. The retrieval of ²⁴⁷ the scattering signal at 3% of the maximum laser intensity means that the lower detection ²⁴⁸ limit for coating-thickness analysis is substantially higher than for delay-time analysis.

 The SP2 was operated and calibrated as described in [*[Laborde et al.](#page-31-4)*, [2012a\]](#page-31-4). Briefly, mass-selected rBC particles representative of atmospheric or diesel rBC (Alfa Aeser Inc., FS, Lot #FS12S011) were used to calibrate the peak incandescence signal and polystyrene 252 latex (PSL) sphere standards of diameter 269 nm were used to calibrate the partial scatter-ing cross section measurements.

²⁵⁴ **2.3 Aerosol Mass Spectrometer (AMS)**

 Particles are introduced into the High-ResolutionAerosol Mass Spectrometer (AMS) by an aerodynamic lens [*[Liu et al.](#page-32-5)*, [2007\]](#page-32-5) which has its maximum transmission efficiency between about 100 nm to 600 nm in free-molecular aerodynamic diameter $(d_{fm,a})$. Note that this aerodynamic diameter is weighted by particle density [*[Kulkarni et al.](#page-31-6)*, [2011\]](#page-31-6), ²⁵⁹ so that $d_{\text{fm,a}} = 100 \text{ nm}$ corresponds to a volume-equivalent diameter of 69 nm for the HFO PM density reported below. Non-refractory material in these particles is then vapor- $_{261}$ ized by an 873 K porous-tungsten cone before electron ionization. The resulting ions are analyzed in a high-resolution time-of-flight mass spectrometer [*[DeCarlo et al.](#page-29-7)*, [2006\]](#page-29-7). Note that we have used the term "sulfate" below rather than "sulfuric acid" (the likely form of sulfate, as discussed below) because sulfate and nitrate are the chemical components mea- sured by AMS after electron ionization. For a more detailed discussion of the AMS re-sults of this study, see *[Corbin et al.](#page-29-5)* [\[2017\]](#page-29-5).

²⁶⁷ **2.4 Optical measurements**

²⁶⁸ The aerosol absorption coefficient was measured by the extinction-minus-scattering ²⁶⁹ [t](#page-33-7)echnique, using a Cavity Attenuation Phase Shift PM SSA monitor [CAPS PMssa, *[Onasch](#page-33-7)* ₂₇₀ *[et al.](#page-33-7)*, [2015\]](#page-33-7) operating at 780 nm. This instrument provides a calibration-free measure of

extinction (manufacturer-estimated accuracy: 5%). The extinction measurement occurs in an optical cell surrounded by an integrating-sphere nephelometer. This nephelometer was calibrated using the extinction measurement, as described in *[Onasch et al.](#page-33-7)* [\[2015\]](#page-33-7), using nm diameter polystyrene latex spheres and size-selected NH_4NO_3 . We also performed 275 a calibration at the Paul Scherrer Institute using pure N_2 and CO_2 . Overall, eight calibra-₂₇₆ tions over 13 months using all of these approaches showed a relative standard deviation of only 10%, indicating very stable instrument performance. Data were manually inspected to ensure that baseline values of extinction and scattering (obtained periodically on filtered samples) were stable. Note that during engine measurements, the CAPS PMssa was oc- casionally diluted by a factor of 5.1 owing to flow limitations in some experiments (not presented here), which was tested and found to have no effect on the resulting data.

 Some data were also acquired using CAPS PMssa instruments operating at 450 nm and 630 nm. These data are not presented here due to issues with data quality, possibly due to the difficulty of obtaining a valid baseline measurement when variable amounts of absorbing gases with strong surface interactions, such as NO_2 , are present.

²⁸⁶ An aethalometer (model AE33, Aerosol d.o.o., Slovenia) was also deployed dur-²⁸⁷ ing these experiments. The AE33 measures the light attenuation through a PM deposit 288 on a filter at seven wavelengths $λ$. The AE33 converts filter-deposit attenuation coeffi-

cients to aerosol light absorption coefficients by applying a conversion factor (known a cients to aerosol light absorption coefficients by applying a conversion factor (known as ²⁹⁰ the "C value") as detailed in *[Drinovec et al.](#page-29-8)* [\[2015\]](#page-29-8) and discussed further in Section S3, ²⁹¹ where the calibration factor is evaluated for our samples at 780 nm. To avoid issues due ²⁹² to pressure changes in the sampling lines due to changes in engine conditions, our AE33 ²⁹³ data was reanalyzed with a constant compensation parameter [defined in *[Drinovec et al.](#page-29-8)*, 2015].

 Three filter samples were also measured with the Multi-Wavelength Absorbance An- alyzer [MWAA; *[Massabò et al.](#page-32-6)*, [2013,](#page-32-6) [2015\]](#page-33-8) with the goal of evaluating the performance of the CAPS PMssa. In particular, the wavelength dependence of the aethalometer calibra- tion could be confirmed as negligible using the MWAA data. This is evident in Fig. [6,](#page-26-0) as discussed further in the supplement.

³⁰⁰ **2.5 Optical calculations**

³⁰¹ Using the AE33 data, the AAE was calculated as the two-wavelength AAE:

$$
AAE(\lambda_1, \lambda_2) = -\frac{\ln (b_{ATN, \lambda_1}/b_{ATN, \lambda_2})}{\ln (\lambda_1/\lambda_2)}
$$
(1)

 302 where the subscripts 1 and 2 indicate two different measurement wavelengths, λ is a 303 measurement wavelength, and $b_{\text{ATN},\lambda}$ is an AE33-measured attenuation coefficient.

- ³⁰⁴ Using the CAPS PMssa data, $b_{abn,\lambda}$ was calculated at λ =780 nm by subtracting
- ³⁰⁵ the measured scattering from the measured extinction coefficient. By combining the CAPS
- ³⁰⁶ PMssa $b_{abn,780}$ with the measured AAEs, we also calculated $b_{abn,\lambda}$ at $\lambda = (370, 470, 520, 590, 660)$ nm.

³⁰⁷ This calculation is numerically equivalent to calibrating the aethalometer using the CAPS

³⁰⁸ PMssa (full details are given in the supplement). For some samples, CAPS PMssa data

- ³⁰⁹ were not available for this calibration, and the median *C*-value from the other calibrations
- 310 was applied.

 311 The MAC of a species *X* at a wavelength λ , MAC_{X, λ}, was calculated as

$$
MAC_{X,\lambda} = b_{abn,X,\lambda} / C_X,
$$
 (2)

 312 where $b_{abn,X,\lambda}$ is the aerosol absorption coefficient corresponding to absorption

³¹³ by species *X*. In general, brC is considered to absorb negligibly at longer wavelengths

³¹⁴ [*[Laskin et al.](#page-31-7)*, [2015\]](#page-31-7), so that we attribute all absorption at $\lambda \ge 780$ nm to BC ($b_{abn,BC,\lambda} \approx b_{abn,\lambda}$) and the equation becomes

 $b_{abn,\lambda}$) and the equation becomes

$$
MAC_{rBC,780} = b_{abn,BC,780} / C_{rBC} \approx b_{abn,780} / C_{rBC},
$$
 (3)

³¹⁶ for BC, with C_{rBC} measured by the SP2, corrected for the mass fraction outside of 317 the SP2 sensitivity range (Section [3.1.1\)](#page-11-0).

³¹⁸ Light absorption was attributed to brC when the measured absorption exceeded that 319 predicted by extrapolating measurements at longer wavelengths ($\lambda \ge 780 \text{ nm}$) to shorter
320 ones ($\lambda < 660 \text{ nm}$) using an AAE= 1. That is, brC absorption was defined by ones ($\lambda \le 660 \text{ nm}$) using an AAE= 1. That is, brC absorption was defined by

$$
b_{\text{abn},\text{brC},\lambda} = b_{\text{abn},\lambda} - b_{\text{abn},\text{BC},880} \left(\frac{\lambda}{880 \text{ nm}}\right)^{-\text{AAE},\text{BC}},\tag{4}
$$

where λ and 880 nm represent two measurement wavelengths in nm and the sub-
tracted quantity is the AAE-based estimate of BC absorption at λ , using AAE_{BC} = 1 ex tracted quantity is the AAE-based estimate of BC absorption at λ , using AAE_{BC}= 1 except 323 during sensitivity tests (see below). Eq. [4](#page-8-0) results in a positive value of $b_{abn,bnC,\lambda}$ when $AAE_{PM} > 1$ and zero $b_{abn,bn}C_{A}$ when $AAE_{PM} = 1$. In no case was $AAE_{PM} < 1$. While ³²⁵ both *^b*abn,λ and the AAE may be affected by substantial internal mixing [*[Lack and Cappa](#page-31-8)*, ³²⁶ [2010\]](#page-31-8), we expect this effect to be negligible in our data because substantial internal mix- 327 ing of brC and BC was ruled out by our measurements (Section [3.1.2\)](#page-13-0). Eq. [4](#page-8-0) was employed to calculate $b_{abn,bn}$ _c, λ at $\lambda = \{370, 470, 520, 590, 660\}$ nm, corresponding to each wavelength shorter than the reference (CAPS PMssa) $\lambda = 780$ nm. wavelength shorter than the reference (CAPS PMssa) $\lambda = 780$ nm.

330 As shown in Eq. [4,](#page-8-0) we generally used $\lambda_2 = 880$ nm as the reference wavelength. A 331 sensitivity test using $\lambda_2 = 950$ nm changed the calculated brC absorption by HFO by only \approx 2% on average. (HFO was the only fuel with AAE_{PM} substantially higher than unity.) ³³³ Uncertainties in the use of Eq. [4](#page-8-0) are discussed in Section [2.7.](#page-10-0)

 334 The MAC of brC was calculated using Eq. [2](#page-8-1) with $b_{abn,bn \in \lambda}$ and C_{OM} substituted for $b_{abn, X, \lambda}$ and C_X , respectively:

$$
MAC_{OM, \lambda} = b_{abn, brC, \lambda} / C_{OM},
$$
 (5)

336 where C_{OM} is the OM mass concentration measured by the AMS.

³³⁷ **2.6 Mie model**

³³⁸ A Mie model was constructed and fitted to the brC absorption and OM mass con- 339 centration measurements to obtain imaginary refractive indices of OM, k_{OM} . The model 340 structure, inputs and outputs are depicted in Fig. [2](#page-23-1) and described in the following.

³⁴¹ As described below, our measurements indicated that these ship-engine emissions ³⁴² can be appropriately modelled as an external mixture of two modes, a larger rBC mode 343 and a smaller nucleation mode (internally mixed OM and sulfate). In estimating $b_{abn,bn}C_{\lambda}$ ³⁴⁴ from Eq. [4,](#page-8-0) we have already subtracted light absorption by BC from the total, and must 345 only model the brC absorption when fitting the Mie model.

 The model therefore consisted of a single lognormal distribution of internally-mixed sulfates and OM; the reported k_{OM} values were obtained by minimizing the difference between the measured absorption and the Mie-predicted absorption of this distribution (each reported *k*OM was fitted independently of the others). The Mie calculations required several additional pieces of information, which were constrained as follows. The GMD was obtained from bimodal lognormal fits to the FMPS measurements and a GSD of 1.65 was obtained from the SMPS data as described in Section [2.1.](#page-3-0) The respective fractions of OM and sulfate in the particles were obtained from the AMS data. The particle mate- rial density was obtained from the DMA–APM measurements described above. For HFO 355 and MGO PM, the respectively measured material densities were 1460 ± 50 kg m⁻³ and 356 900 \pm 40 kg m⁻³. The higher density for HFO PM is due to the presence of sulfate. The density measurements of MGO PM were also used to estimate the density of the DF PM, which was not measured. The measured densities are in good agreement with the density estimated from the mass spectra of these samples [*[Kuwata et al.](#page-31-9)*, [2012\]](#page-31-9), as described in the supplement.

361 The real refractive index of OM, n_{OM} was taken as $n = 1.5$ [*[Lu et al.](#page-32-7)*, [2015\]](#page-32-7); the ³⁶² fitted results did not change when varying this value as $n_{OM} \pm 0.1$. The corresponding $n_{SO²}$ was taken as 1.35 (sulfuric acid). The overall refractive index of the particles was ³⁶³ *n*_{SO}²⁻ was taken as 1.35 (sulfuric acid). The overall refractive index of the particles was ³⁶⁴ obtained by combining *n*_{OM} and *n*_{SO²⁻} according to their AMS-measured volume ratios, ³⁶⁵ following the volume-weighted linear mixing rule validated by *[Abo Riziq et al.](#page-29-9)* [\[2007\]](#page-29-9). ³⁶⁶ Sensitivity tests indicated that the most sensitive input parameter in this model was the 367 GMD: a large change in GMD of 20 nm corresponded to a change of only ± 0.001 in ³⁶⁸ *k*OM (*k*OM was greater than 0.01 for all HFO samples, as shown in Fig. [8](#page-27-0) below). Varying ³⁶⁹ other parameters by large amounts (relative to their anticipated uncertainties) generally led

 $10³⁷⁰$ to < 10 % change in k_{OM} . We note that the PM density used here was well-constrained by the APM measurements. the APM measurements.

³⁷² Overall uncertainties in k_{OM} were calculated by propagating the numerical uncer-373 tainties given below in a Monte Carlo calculation.

³⁷⁴ **2.7 Statistical uncertainties**

³⁷⁵ Unless specified otherwise in this section, uncertainties are statistical imprecisions and are reported as the standard error of the mean $(\sigma_{\bar{x}} = SD/\sqrt{(N)}, SD =$ standard devi-³⁷⁷ ation) with the averaging periods defined by the 30 minute filter periods mentioned above. For derived quantities, uncertainties were propagated from these $\sigma_{\bar{x}}$.

³⁷⁹ In the special cases where uncertainties other than statistical noise were known to be ³⁸⁰ significant, those uncertainties were added in quadrature (i.e., possible covariance between ³⁸¹ these different uncertainties was not addressed):

³⁸² • For the CAPS PMssa, uncertainties in b_{abn} were propagated from an estimated 5 % α ³⁸³ accuracy in b_{ext} [as estimated by *[Onasch et al.](#page-33-7)*, [2015\]](#page-33-7), 10% accuracy in scatter- δ ₃₈₄ ing calibration (based on the SD of multiple calibrations), $\pm 1 \text{ Mm}^{-1}$ baselining ³⁸⁵ accuracy and 5 % accuracy in dilution correction. The fraction of scattered light ³⁸⁶ not collected by the CAPS PMssa integrating sphere (the truncation error) must be ³⁸⁷ corrected for if the phase function of measured particles is substantially different ³⁸⁸ to that of the calibration particles. This is not the case if both measured and cali-³⁸⁹ bration particles are small relative to the 780 nm wavelength of the CAPS *[Onasch](#page-33-7)* ³⁹⁰ *[et al.](#page-33-7)* [\[2015\]](#page-33-7), which was generally true. Any potential complications introduced by ³⁹¹ the larger BC mode in HFO PM are unlikely, considering the fact that the HFO ³⁹² data were not outliers in the optical analyses described below. We note that the calculated absorption is not highly sensitive to truncation errors at low SSA values $394 \left(\sim 0.7. \right)$

 • For the SP2, the relatively low number of rBC particles measured in each sampling period yielded substantial statistical imprecision, which is the major source of re- ported uncertainty. Calibration-related uncertainties were not propagated into the reported uncertainties; the precision of our calibration data was much less than the just-mentioned statistical imprecision. An unknown source of uncertainty is cal- ibration bias related to differences between calibrant and sample. To the extent that marine-engine rBC is similar to road-vehicle rBC, the latter bias is estimated as < 14% [*[Laborde et al.](#page-31-5)*, [2012b\]](#page-31-5).

• For the AMS, a measurement precision of $0.3 \mu g m^{-3}$ was estimated from mea- surements of filtered air and combined with the statistical imprecision of the mean. 405 Measured concentrations may have been biased low by ~ 10 to 35 %, for smaller particles, due to the small measured particle sizes (e.g. ∼ 60 nm mode diameter of volume distributions). This size-dependent bias was discussed by *[Liu et al.](#page-32-5)* [\[2007\]](#page-32-5). Other known biases [*[Jimenez et al.](#page-30-6)*, [2016\]](#page-30-6) include variability in the electron ion-

uncertainties of $b_{abn,bn}C_{\lambda}$.

⁴¹⁶ • In calculating k_{OM} , estimated imprecisions of 5 nm in GMD, 0.1 in GSD, and 4 % ⁴¹⁷ in density (from the SD of multiple measurements) were considered as Gaussian and propagated by Monte Carlo calculation.

3 Results and discussion

3.1 Engine emissions

 The emission factors relative to engine power (in units of mg / kWh) for OM, sul- fates, rBC and nitrates from this engine were similar to those reported by *[Mueller et al.](#page-33-1)* [\[2015\]](#page-33-1) for this engine. For all three fuels, OM emissions dominated the total PM mass ⁴²⁴ with little sensitivity to engine load. For HFO, OM contributed 58 % of the $PM_{2.5}$ mass at 425 50 % load (because sulfate contributed 25 %) and > 82 % of the PM_{2.5} mass below 50 % $\frac{426}{426}$ load. Notably, when normalized to $CO₂$ rather than engine power, the variability in emis-⁴²⁷ sion factors for rBC was substantially reduced. Section S1 provides more detail on these emission factors.

3.1.1 Total & BC size distributions

 Particle number and volume size distributions, as measured by SMPS as a func- tion of mobility diameter d_{mob} , are shown in Fig. [3.](#page-24-0) Also shown in these figures are the ⁴³² number and volume distributions of rBC cores, as measured by SP2 as a function of d_{rBC} . These size distributions are representative averages of selected measurement periods, where ⁴³⁴ both SMPS and SP2 data were available.

 The size distributions of the diesel fuel (DF) and marine gas oil (MGO) PM, shown in Fig. [3,](#page-24-0) were comparable to that of typical diesel engines [*[Burtscher](#page-29-10)*, [2005\]](#page-29-10), in that two separate modes are identifiable, a volatile ultrafine mode (at about 60 nm d_{mob} in the totalvolume distribution) and a soot mode (lognormal at about 120 nm d_{rBC} in the rBC-volume distribution) were observed. AMS measurements of the fresh and thermodenuded aerosol indicated that the ultrafine mode for these distillate fuels consisted mainly of organics and, ⁴⁴¹ for MGO, sulfates. The rBC volume distribution indicated that a fraction of the MGO and ⁴⁴² DF particles was above or below the quantification range of the SP2, corresponding to an 443 unmeasured mass fraction of $\lt 3\%$ (above) and $\lt 23\%$ (below), respectively. The rBC mass concentrations reported from these SP2 data have been corrected for this missing fraction, as discussed further at the end of this subsection.

 In considering Fig. [3](#page-24-0) relative to previous studies [*[Oeder et al.](#page-33-0)*, [2015;](#page-33-0) *[Mueller et al.](#page-33-1)*, [2015;](#page-33-1) *[Streibel et al.](#page-35-2)*, [2017\]](#page-35-2), it is important to realize that the SP2-measured d_{rBC} is a vol-⁴⁴⁸ ume equivalent diameter, which is by definition smaller than d_{mob} for soot particles of both open and compact morphologies [*[Kulkarni et al.](#page-31-6)*, [2011\]](#page-31-6). For the two size distri-450 butions in Fig. [3](#page-24-0) to overlap, a size-independent shape factor of \sim 2.2 is required. This value is within the range of values (1.5–3) recommended by *Sorensen* [2011] for uncoa value is within the range of values (1.5–3) recommended by *[Sorensen](#page-35-7)* [\[2011\]](#page-35-7) for uncoated, fractal-like BC aggregates, and corresponds to ∼ 50 primary particles in a typical BC ag- gregate, if the monomer diameter is 30 nm [this monomer diameter is from *[Oeder et al.](#page-33-0)*, [2015\]](#page-33-0). Thus, our measured size distributions are consistent with the evidence presented in 455 Section [3.1.2](#page-13-0) to demonstrate that rBC was externally mixed.

 The size distribution of the residual fuel, HFO, also showed an ultrafine mode con- sisting of sulfates and organics with a mode diameter of about 90 nm. This mode dom-⁴⁵⁸ inated the overall particle number and had a mean $(\pm SD)$ effective density of 1460 \pm 50 kg m⁻³. The effective density was considerably lower for MGO, at 890 \pm 30 kg m⁻³ (data shown in Fig. [3\)](#page-24-0). The effective densities for DF PM was not measured. Since the [p](#page-34-6)articles consisted of organics and sulfate (see also Section [3.1.2\)](#page-13-0) they were liquid [*[Se-](#page-34-6) [infeld and Pandis](#page-34-6)*, [2012\]](#page-34-6), and these effective densities correspond directly to the material density of this PM mode. We could not obtain effective density measurements of rBC-containing particles due to their relatively low number concentrations.

 The rBC-core distribution showed a much broader volume distribution than is ex- [p](#page-34-7)ected for vehicular [*[Burtscher](#page-29-10)*, [2005;](#page-29-10) *[Laborde et al.](#page-31-10)*, [2012c\]](#page-31-10), biomass burning [*[Schwarz](#page-34-7) [et al.](#page-34-7)*, [2008;](#page-34-7) *[Laborde et al.](#page-31-10)*, [2012c;](#page-31-10) *[Wang et al.](#page-35-8)*, [2016\]](#page-35-8), or aircraft [*[Lobo et al.](#page-32-8)*, [2015\]](#page-32-8) emis- sions. Two studies of Chinese air pollution have also reported large rBC modes of 690 nm [*[Huang et al.](#page-30-8)*, [2011\]](#page-30-8) and 610 nm [*[Wu et al.](#page-35-9)*, [2017\]](#page-35-9), which may be linked to HFO-combustion pollution. In HFO exhaust from a marine diesel engine, *[Kasper et al.](#page-30-9)* [\[2007\]](#page-30-9) also observed a large, non-volatile (at 400°C) mode, but did not measure its chemical composition. Our measurements show that at least part of this mode consists of rBC. It likely comprised the char particles defined in the introduction. Future studies should further explore the proper-ties of this large rBC mode.

 The SP2 did not measure the tail of the hypothesized char mode (larger rBC mode). In the laboratory, *[Linak et al.](#page-32-4)* [\[2000\]](#page-32-4) found that char particles produced from HFO had a typical diameter of 1 μ m, which may correspond to the significant mass fraction of PM
that has been identified in the coarse mode of marine engines [*Fridell et al.*, 2008]. The that has been identified in the coarse mode of marine engines [*[Fridell et al.](#page-30-10)*, [2008\]](#page-30-10). There- fore, we initially hypothesized that the SP2 may not have measured all BC mass in the aerosol. However, we did not find supporting evidence for this hypothesis, as discussed in the following paragraphs.

 A simple test of this hypothesis may be performed using the SP2 data directly. At the upper limit of single-particle mass, single particles are measured as saturated signals in the SP2. For the HFO size distribution presented in Fig. [3](#page-24-0) only 5 of 350,000 parti- cles caused saturated signals. Therefore, the majority of the PM passing through our sam-⁴⁸⁶ pling system (which included a $PM_{2.5}$ cyclone) and reaching the SP2 was measured. This

 conclusion is corroborated by the excellent correlation between CAPS PMssa and SP2 (Fig. S1 and S2), for which data were available only for selected experiments.

 Another test of this hypothesis was performed using Mie calculations of the AAE(880,950) for BC and for cenospheres. We chose this wavelength pair so that absorption by canoni- cal brC would be negligible [*[Laskin et al.](#page-31-7)*, [2015\]](#page-31-7). Cenospheres are core-shell BC particles with air cores and have been observed in HFO PM [e.g., *[Chen et al.](#page-29-4)*, [2005\]](#page-29-4). They may be modelled accurately with core-shell Mie theory [*[Huang et al.](#page-30-11)*, [2012\]](#page-30-11), which we have done ⁴⁹⁴ in Fig. S3. The cenospheres were modelled with a ratio of inner to outer diameters of 0.8 (the abscissa of Fig. S3 is the outer diameter).

⁴⁹⁶ The Mie calculations (Fig. S3) show that in the Rayleigh regime (diameter $d <$
⁴⁹⁷ 30 nm) the predicted AAE_{PC} \approx 1 does not depend on the model as expected. Outside 30 nm) the predicted AAE $_{BC} \approx 1$ does not depend on the model, as expected. Outside of the Rayleigh regime (*d* > 400 nm), the model becomes important; in our calculations the
AAE_{BC} \approx 0. The calculated AAE_{BC} \approx 0 is much smaller than our observed AAE(880.950 AAE_{BC} \approx 0. The calculated AAE_{BC} \approx 0 is much smaller than our observed AAE(880,950) $500 \geq 1$ (also shown in the figure), suggesting that the majority of BC absorption and therefore BC mass did not reside in unmeasured char particles.

 Overall, the HFO rBC size distribution could be described by a bimodal lognormal fit with geometric mean d_{rBC} and standard deviations (GMD and GSD) of 150 and 640 nm, and 1.0 and 0.56, for the two modes. These lognormal fits imply that 44% of the total mass was likely above (versus 2% below) the particle size range detectable by the SP2 and were used to correct the total reported rBC mass. The corresponding SP2-mass correction factor was 1.46 ± 0.15 . This uncertainty represents inter-experiment variability and was propagated into the overall uncertainty. For unimodal lognormal fits to the diesel 509 and MGO data, the correction factors were 1.247 ± 0.05 and 1.228 ± 0.05 , respectively. These latter correction factors account for particles which were too small to be measured by the SP2.

3.1.2 External mixing of rBC and OM/sulfate

 Fig. [4A](#page-25-0) and [4B](#page-25-0) investigate the coating thickness on BC particles by analyzing the time-resolved SP2 scattering signals in two different ways. Fig. [4A](#page-25-0) shows that the num- ber fraction of particles with evidence of 'thick coatings' (where the peak scattering sig- nal occurs before particle incandescence, indicating the loss of a large volume of coating material due to heating) was negligible except for smaller DF particles. (Since no brC ab- sorption was observed for DF, this exception does not influence any of the conclusions made in this study.) The remainder, 'moderate or no coating'', describes particles where ₅₂₀ peak scattering occurred simultaneously with incandescence. This delay-time analysis pro- vides a relatively coarse picture of the mixing state of BC; more quantitative information is provided by the coating-thickness analysis.

 Fig. [4B](#page-25-0) shows the results of the SP2 coating-thickness analysis. For this analysis, only BC particles with spherical-equivalent diameter 240 \pm 20 nm were included, due ₅₂₅ to the need to optimize both counting statistics and signal-to-noise in the SP2 detectors.

₅₂₆ Consistent with the delay-time analysis, the coating-thickness results show that the mean coating thickness for HFO particles was negligible, whereas small amounts of coatings may have been present for MGO and DF. Negative coating thicknesses reflect particle-to- particle errors (biases) in the scattering measurements, which average to zero. Note that an order of magnitude more data were available for the HFO analysis than for the other fuels analysis. In all cases, the analysis indicates that the coatings were not thick enough to influence the optical properties of the BC [*[Liu et al.](#page-32-9)*, [2017\]](#page-32-9).

 Further analysis of independent size-resolved composition measurements for HFO [*[Corbin et al.](#page-29-5)*, [2017\]](#page-29-5) support this picture. Ash (metal oxides and/or sulfates) is also ex- pected in the HFO PM, as discussed in *[Corbin et al.](#page-29-5)* [\[2017\]](#page-29-5). However, with the measure- ments presented here, we can only infer that ash particles did not make up the majority of ₅₃₇ the total particle mass for either rBC or OM/sulfate particles.

 The effective density measurements for the mobility size range $60 - 120$ nm shown in Fig. [3](#page-24-0) also support our conclusion that the smaller mode in that figure consisted of externally-mixed OM (and sulfate, in the case of HFO) . The effective densities (calculated as $6m_p/[{\pi}d_{\text{mobility}}^3]$, where m_p is the APM-measured single-particle mass and d_{mobility} is ₅₄₂ the DMA-selected diameter, as further described in the supplement) are shown by the inset circles in Fig. [3.](#page-24-0) For MGO and DF, these densities should correspond to the mate- $_{544}$ rial density of the OM in those fuels. Indeed, the measurements are in close agreement with the densities predicted from a published empirical relationship between the elemental composition and the density of OM [*[Kuwata et al.](#page-31-9)*, [2012\]](#page-31-9). That parameterization showed a low bias of just 5% (standard deviation 2%) relative to the DMA-APM-CPC measure-ments.

3.1.3 Relationship of BC light absorption with SSA780 nm and engine load

 BC light absorption was characterized according to the aerosol light absorption co-⁵⁵¹ efficient measured at 780 nm, $b_{abn,780nm}$, which was measured by the CAPS PMssa using ₅₅₂ the extinction-minus-scattering technique. The measured extinction and scattering coef- ficients also provided the aerosol single-scattering albedo ($SSA_{780 \text{ nm}}$; Fig. [5A](#page-25-1)). All three fuels showed a range of $SSA_{780 \text{ nm}}$ (from 0.5 to 0.9). The lower SSAs correspond to con- ditions where the extinction-minus-scattering method results in a more accurate calculation of $b_{abn,780nm}$.

 The ratio of CAPS $b_{abn,780nm}$ to SP2 C_{rBC} gives the MAC_{BC,780nm} (Eq. [3\)](#page-8-2), as shown 558 in Fig. [5B](#page-25-1). The MACBC,780nm did not vary substantially between engine loads or fuel $_{559}$ types. The $SSA_{780 \text{ nm}}$ did vary substantially, but without a clear dependence on engine load, fuel type, or indeed any other parameter we evaluated, including the GMD (geomet-⁵⁶¹ ric mean diameter of the FMPS) and the ratio BC:non-refractory-PM. These quantities were explored because it was expected that the MAC_{BC,780nm} and SSA_{780nm} would vary as a function of aerosol composition (for example, the mass fraction of BC), size (since scattering depends strongly on particle size), or mixing state (since internal mixing of BC with other material may result in absorption enhancements). Note that although the

⁵⁶⁶ accuracy of the CAPS-estimated $b_{\text{abn},780\text{nm}}$ is expected to decrease with increasing SSA, our data indicate that this accuracy was not a limiting factor: there was no statistically- $_{568}$ significant dependence of the $MAC_{BC,780nm}$ on the SSA.

 The lack of any such observed relationship indicates that the corresponding phenom- ena were not the main cause of variations in the $SSA_{780 \text{ nm}}$. Moreover, the fact that varia- tions in SSA_{780 nm} did not correspond to variations in MAC_{BC,780 nm} is consistent with the aerosol containing a variable fraction of externally-mixed BC. We therefore hypothesize that variable amounts of externally-mixed ash particles were the main cause of the vari- ability in the SSA 780 nm . Such ash particles were identified by electron microscopy in these ₅₇₅ samples [*[Corbin et al.](#page-29-5)*, [2017\]](#page-29-5), but more data would be needed to quantitatively evaluate this hypothesis.

 The shading in Fig. [5B](#page-25-1) shows the result of an uncertainty-weighted orthogonal-578 distance regression fit to a plot of $b_{\text{abn},780\text{nm}}$ vs C_{rBC} (the shading shows the fit $\pm 95\%$ 579 confidence interval; Fig. S1). As there was no apparent difference in MAC_{BC,780 nm} be-580 tween different fuels, all data were fit simultaneously to yield $MAC_{BC,780 \text{ nm}} = 7.8 \pm 1.8 \text{ m}^2 \text{ g}^{-1}$. Assuming an AAE_{BC} of 1, this corresponds to a MAC_{BC} 550 nm of 11.1 \pm $1.8 \text{ m}^2 \text{ g}^{-1}$. Assuming an AAE_{BC} of 1, this corresponds to a MAC_{rBC,550 nm} of 11.1 \pm 2.6, a factor 1.48 larger than the widely-accepted MAC_{BC,550 nm}= $7.5 \text{ m}^2 \text{ g}^{-1}$ recommended by *[Bond and Bergstrom](#page-29-6)* [\[2006\]](#page-29-6) for uncoated BC.

 An increase in MAC due to so-called lensing effects may occur if rBC is internally- mixed with substantial amounts of non-absorbing material [*[Liu et al.](#page-32-9)*, [2017\]](#page-32-9). However, our SP2 measurements indicate that much more internal mixing than observed would be necessary to explain a 48% lensing enhancement (Fig. [4\)](#page-25-0). The calculated absorption en- hancement for a 240 nm rBC core is shown on the upper axis of Fig. [4,](#page-25-0) using a Mie ap- proximation. For HFO, a negligible enhancement is predicted. For MGO and DF, a max- $_{590}$ imum absorption enhancement of 10–20 % is predicted. Further calculations verified that these enhancements change negligibly when the soot aggregates are represented as non- interacting monomer spheres, instead of volume-equivalent spheres [*[Liu et al.](#page-32-10)*, [2015\]](#page-32-10). Con- sidering the uncertainties associated with the coating-thickness and MAC data, we con-₅₉₄ sider that the discrepancy between the calculated and apparent absorption enhancements is probably negligible for MGO and DF. For HFO, where our confidence is enhanced by a larger number of analyzed particles, we find that lensing does not explain the why the $MAC_{rBC,780\,\text{nm}}$ was higher than expected. Further work is needed to clarify why this HFO 598 MAC_{rBC,780 nm} was higher than previously-reported values.

 Some difference in MAC_{rBC,780 nm} for HFO compared to the distillate fuels may have ₆₀₀ been anticipated due to the presence of a large rBC mode in HFO PM (Section [3.1.1\)](#page-11-0). ⁶⁰¹ This was not observed. One may therefore hypothesize that all rBC particles consisted of soot aggregates, since the MAC_{rBC},780 nm of a soot aggregate depends primarily on the size of its monomers [*[Sorensen](#page-35-10)*, [2001\]](#page-35-10) and not on its overall size. However, as measure- ments on char particles are lacking in the literature, it is not clear that the $MAC_{rBC,780 nm}$ (and other optical properties) of small char particles would be sufficiently different from soot to be seen in Figs. [5B](#page-25-1) and S1. Regardless of the physical interpretation, our data

⁶⁰⁷ indicate that the anomalous size distribution of HFO PM did not lead to an anomalous

 $MAC_{rBC,780 \text{ nm}}$. We note, however, that SP2 measurements were not available for the low-

⁶⁰⁹ est engine load test points, for which brC absorption appeared to be anomalously high.

⁶¹⁰ **3.2 Brown carbon**

 ϵ ₆₁₁ In this section, we discuss the wavelength-dependence and intensive properties of ⁶¹² brown carbon (brC). The focus is on HFO PM, since brC was negligible in the other fu-⁶¹³ els.

We reiterate here that these calculations are numerically independent of the SP2 data 615 discussed in Section [3.1.1,](#page-11-0) as shown in the schematic diagram Fig. [2.](#page-23-1) Uncertainties re-616 lated to the larger mode of rBC particles measured in HFO PM by SP2 are therefore ir-617 relevant to the following results. In particular, we used only the $b_{abn,780nm}$ measured by 618 CAPS PMssa and the AAE measured by the AE33 to calculate k_{OM} . The assumptions are 619 that $AAE_{BC} = 1$ and that only BC absorbs at $\lambda > 700$ nm.

⁶²⁰ *3.2.1 Wavelength dependence of light absorption*

 ϵ_{21} Fig. [6](#page-26-0) shows AE33 absorption spectra for all engine test points. The data are nor-⁶²² malized to 880 nm to allow different concentrations to be compared. The distillate fuels 623 (DF and MGO) show similar trends and are consistent with an AAE of 1 ± 0.1 down to 470 nm. At 370 nm, there is some evidence of a minor contribution (< 20% of b_{abn} at 370 nm) of brC to light absorption for the distillate fuels 370 nm) of brC to light absorption for the distillate fuels.

For HFO, the data are generally consistent with an AAE(370–950) of 1.7 ± 0.2 .

The shaded region in Fig. 6 encompasses this range, showing that this AAE includes t The shaded region in Fig. [6](#page-26-0) encompasses this range, showing that this AAE includes the majority of the data points while slightly underpredicting absorption at $\lambda < 500$ nm. We
hypothesized that the scatter in Fig 6 may be related to a varying mass fraction of BC hypothesized that the scatter in Fig [6](#page-26-0) may be related to a varying mass fraction of BC 630 (relative to OM and sulfate) but found this this was not the case (Fig. S4).

 F ig. [6](#page-26-0) also includes MWAA measurements of light absorption of the filter samples. ₆₃₂ These spectral measurements provide additional confidence in the AE33 data, since the ⁶³³ MWAA directly measures and corrects for light scattering from the filter sample. The 634 MWAA data represent only three HFO filter samples due to limited sample availability, ⁶³⁵ but generally corroborate the AE33-measured AAEs.

⁶³⁶ *3.2.2 MACOM,370*

 F ₆₃₇ Fig. [7C](#page-27-1) shows brC absorption at 370 nm, per mass of OM (MAC_{OM,370}) for all fu-⁶³⁸ els as a function of engine load. For cases in which the median AE33 *C* value was used, ⁶³⁹ rather than a specifically-calibrated *C*-value, the data are shown with open rather than ⁶⁴⁰ filled symbols. The upper panel, Fig. [7A](#page-27-1), illustrates the AE33-measured AAEs from which ⁶⁴¹ these MACs have been derived.

 F ig. [7C](#page-27-1) shows that the MAC_{OM,370} increased by an order of magnitude at the low-⁶⁴³ est engine load for HFO. This increase did not correspond to a change in the overall OM $_{644}$ emission factor EF_{OM}, as seen by comparison to Fig. [7B](#page-27-1), nor to a change in the sampled 645 OM concentrations (not shown). Therefore, this change must be due to a difference in the ⁶⁴⁶ chemical composition of the OM. For this reason, we hypothesized that the less-volatile 647 thermal–optical OC fractions would correlate with $MAC_{OM,370}$, but found that the cor-⁶⁴⁸ relation was poor (Section S5). The relative emission factors of polycyclic aromatic hy-⁶⁴⁹ drocarbons (PAHs), including oxygenated- and nitrogenated PAHs, have been reported to ⁶⁵⁰ increase at low engine loads by *[Sippula et al.](#page-34-8)* [\[2014\]](#page-34-8). Such species have been connected to ⁶⁵¹ the chemical composition of the fuel [*[Rüger et al.](#page-34-9)*, [2015;](#page-34-9) *[Streibel et al.](#page-35-2)*, [2017\]](#page-35-2). The opti-⁶⁵² cal properties are thus not adequately described by the lubricating-oil model proposed by ⁶⁵³ *[Eichler et al.](#page-30-7)* [\[2017\]](#page-30-7), who only discussed engine emissions at 50 % load.

⁶⁵⁴ *3.2.3 Imaginary refractive index of OM, ^kOM*,λ

⁶⁵⁵ Based on the measurements described in the previous sections, a Mie model was ⁶⁵⁶ constructed and fit to the data to obtain the imaginary refractive index of OM for this ⁶⁵⁷ study, as detailed in Section [2.5.](#page-7-1)

 F Figs. [8](#page-27-0) and [9](#page-28-0) show the results of the Mie model fits. Fig. 8 shows $k_{OM,370}$ as a ϵ_{659} function of MAC_{OM,370 nm}. The two quantities are correlated because the main variations ⁶⁶⁰ in model inputs were in OM mass concentration and light absorption, rather than parti-⁶⁶¹ cle size or density. Two extreme outliers have a $k_{OM,370}$ close to unity, which is unusually ₆₆₂ high. We carefully inspected all input data for these measurements but found no errors. ⁶⁶³ The $k_{OM,370}$ for DF and MGO were not significantly different from zero.

 F ig. [9](#page-28-0) depicts the wavelength dependence of k_{OM} by normalizing $k_{OM, \lambda}$ to $k_{OM, 370}$ ⁶⁶⁵ for all data. The grey shading and lines show the behaviour that would be expected for a ⁶⁶⁶ *k*_{OM} AAE of 2, 3, and 4 (upper line, central line, lower line). The plot was filtered to only ϵ_{667} included points where $k_{OM,\lambda} > 0.005$ before normalization. This threshold is based on the Monte-Carlo-estimated uncertainties shown in Fig. [8](#page-27-0) and resulted in the exclusion of all ⁶⁶⁹ MGO and DF data.

 ϵ_{670} Two types of outliers are noticeable in Fig. [9.](#page-28-0) The first, "outlier I", shows an anoma- 671 lously slow decrease in k_{OM} with $λ$. The reason for this exception was not identified. 672 The second, "outlier II", correspond to the two high outliers in Fig. [8.](#page-27-0) The k_{OM} -versus- λ trend for both of these type-II outliers appears reasonable and is similar; compared 674 to the other measurements the slope is more negative than expected. In other words, in ⁶⁷⁵ addition to having higher absolute $k_{OM,370}$ these measurements also had a higher ratio $k_{OM,370}$: $k_{OM,660}$, which is consistent with a true chemical difference in the type-II sam- 677 ples.

4 Discussion

4.1 Expected differences between engines and HFO samples

 In general, the intensive optical properties of marine-engine PM may vary between ⁶⁸¹ studies. In our present study, we have observed similar trends to those observed in the study described by *[Mueller et al.](#page-33-1)* [\[2015\]](#page-33-1), as discussed in detail in Section S5. Further measurements are required to investigate the degree to which engine-to-engine variabil-ity may influence these properties.

 In addition, differences in fuel composition may influence marine-engine PM optical properties. Since the molecular composition of HFO PM has been shown to be compara- ble with the molecular composition of HFO [*[Streibel et al.](#page-35-2)*, [2017\]](#page-35-2), in a manner depending on engine load [*[Sippula et al.](#page-34-8)*, [2014\]](#page-34-8), we hypothesize that the molecular composition of brC is also related to the molecular composition of the fuel. In support of this hypothesis, 690 we note that *[Sippula et al.](#page-34-8)* [\[2014\]](#page-34-8) observed larger mass fractions of PAHs in HFO OM at 691 lower engine loads, similar to our observation of increased MAC_{OM} at lower engine loads. Those authors suggested that the PAHs in HFO PM represent unburned fuel. This picture is consistent with the non-absorbing nature of the distillate-fuel PM samples, since neither MGO nor DF contain substantial levels of PAHs.

 If PAHs in HFO are a major source of brC in HFO PM, then different engines may ₆₉₆ produce brC in a similar manner to one another. Thus, the extrapolation of the MAC_{OM} ⁶⁹⁷ and k_{OM} reported herein to other engines may be justified, especially when using the re-698 lationship of OC3 / OC with MAC_{OM,370} (Fig. S5) to estimate light-absorption by brC in HFO PM. However, since variability in the composition of HFO fuel [*[Stout and Wang](#page-35-11)*, [2016\]](#page-35-11) may lead to variability in the light-absorbing properties of the emissions, additional τ_{01} studies are warranted to constrain fuel-related variability in k_{OM} and MAC_{OM}.

4.2 Summary and implications of brC / OM light absorption

 In the following, we summarize the results on brC absorption and provide an illus-tration of the potential radiative impacts of HFO brC.

 The standard operating load for this engine is 50 %, although engines frequently op- erate at lower loads to save fuel or in the presence of ice [*[Lack and Corbett](#page-31-1)*, [2012\]](#page-31-1). At 50% load, the mean $k_{OM,370nm}$ and corresponding SD were 0.057 \pm 0.027 (*n* = 5). At 25 % load, the corresponding values were 0.034 \pm 0.018 ($n = 6$). Considering the large variabilities, we recommend the combined mean, 0.045 ± 0.025 , for modelling studies. According to Fig. [9,](#page-28-0) $k_{OM, \lambda}$ may then be estimated as

$$
k_{\text{OM},\lambda} = (0.045 \pm 0.025)(\lambda/370 \text{nm})^{-3}, \qquad 370 \text{ nm} < \lambda < 660 \text{ nm}
$$
 (6)

211 At the commonly-used λ of 550 nm, Equation [6](#page-18-0) gives $k_{OM,550nm} = 0.014 \pm 55\%$.

Compared to the *k*OM,550nm values summarized by the review of *[Lu et al.](#page-32-7)* [\[2015\]](#page-32-7), this is a

high value. For example, those authors recommended $k_{OM,550nm} = 0.015 \pm 45\%$ for lig-

ite OM and $k_{OM,550nm} \approx 0.03$ for biomass-combustion OM with 10% BC (a comparable nite OM and $k_{OM,550nm} \approx 0.03$ for biomass-combustion OM with 10 % BC (a comparable 715 fraction to our data). For HFO, *[Lu et al.](#page-32-7)* [\[2015\]](#page-32-7) recommended $k_{OM,550nm} = 0.006 \pm 40\%$, based on the extrapolation of studies of DF emissions to HFO. This is a factor of 2.3 ⁷¹⁷ smaller than our measured value. Considering that we measured negligible brC for DF PM, this extrapolation is not valid and future studies should use Equation [6](#page-18-0) to estimate k_{OM} for HFO.

 The very high OM fraction (approximately 90%) in HFO emissions makes brC light ₇₂₁ absorption especially important for this aerosol in spite of its k_{OM} falling within the range of previously-reported values (at $25-50\%$ engine load). This is demonstrated in Fig. [10,](#page-28-1) where the Simple Forcing Efficiency [*[Chen and Bond](#page-29-11)*, [2010;](#page-29-11) *[Bond and Bergstrom](#page-29-6)*, [2006\]](#page-29-6) has been employed. The figure illustrates the Direct Radiative Forcing (DRF) of a rep- resentative ship-exhaust plume over the planet's surface. A wavelength-dependent mass scattering cross section of 2.3 m² g⁻¹ at 780 nm and scaling exponent of 4 (see supple- ment for details) was used with a backscatter fraction of 0.1. For the 50 % load case (in- puts shown in lower panel), the figure plots the relative absorption of OM and BC (middle panel) and the estimated change in direct radiative forcing (DRF, upper panel) when in- cluding brC absorption. The change in DRF is relative to the base case of non-absorbing OM (equal OM mass, with k_{OM} set to zero). Relative humidity effects have been ne- glected for simplicity. The DRF is highly sensitive to the albedo of the surface below the aerosol; as such, two cases are shown: one for the case of fresh snow or cloud, one for the case of the dark ocean.

 Note that OM absorption from HFO at $λ = 660$ nm is $∼ 20%$ of the total. At this wavelength, brC absorption may have been expected to have dropped to zero [*[Laskin et al.](#page-31-7)*, [2015\]](#page-31-7): HFO brC clearly absorbs farther into the red than expected. At $\lambda = 370$ nm, the contribution of brC to the DRF increases to 18% for the over-snow case. Note that we contribution of brC to the DRF increases to 18% for the over-snow case. Note that we have chosen a conservative example of 50% load, and that MAC_{OM} was much larger at lower loads (Fig. [7\)](#page-27-1). The ability of brC absorption to influence the DRF of HFO-exhaust PM, particularly in high-albedo environments like the Arctic, should motivate modelling studies to assess the overall climate effects of this organic absorption.

5 Conclusions

 We used in-situ measurements of single-particle black carbon mass (via laser-induced incandescence), in-situ aerosol light extinction and scattering coefficients, and filter-based measurements of Ångström absorption exponents, to reach several conclusions on the properties of BC and brC in emissions from an auxiliary-scale research ship engine. One residual fuel (heavy fuel oil; HFO) and two distillate fuels (marine gas oil, MGO, and diesel fuel, DF) were used.

 Overall, light absorption by OM in DF and MGO exhaust was found to be negligi-ble, whereas absorption by OM in HFO PM was very significant. This result is not sur-

and the molecular similarity between HFO-exhaust OM and the fuel [*[Sippula et al.](#page-34-8)*, [2014;](#page-34-8)

[Streibel et al.](#page-35-2), [2017\]](#page-35-2). In more detail, our conclusions are as follows.

- Although the emissions and light-absorbing properties of OM may vary between en- gines, between engine settings, and between HFO samples, our reported k_{OM} and MAC_{OM} values provide a significant improvement over previously-available data.
- **Author contributions**

 JCC performed the initial data analysis and interpretation. HC, MG, IEH, GM, and LD contributed significantly to data interpretation. MZ, JCC and MG operated and cali- brated the SP2. SP and JCC operated and calibrated the CAPS. GJ operated the AE33 and provided its data, LD analyzed the AE33 raw data. AAM, SP, and JCC operated and cal-ibrated the AMS. JCC performed the APM measurements. JO and BS planned and oversaw the experiments. ASHP and RZ initiated the study. JCC wrote the manuscript and all

authors contributed to the writing.

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- 800 615922-BLACARAT. GM and LD are employed by Aerosol d.o.o., where the Aethalome-
- ⁸⁰¹ ter AE33 was developed and is being produced. Emission factor and brC properties data
- 802 are included in the online supplementary material.

Table 1. Common abbreviations and symbols used in the text. Rarely-used abbreviations have been omitted 803

for brevity. 804

Figure 1. Experimental setup. Filter samples were measured after a factor ∼ 12 dilution; other measure-805

Figure 2. Schematic of the BC+brC analysis and Mie model fit performed herein. Symbols and abbreviations are defined in Table [1.](#page-22-0) 807 808

Figure 3. Size distributions of PM from the marine engine operated on (A) HFO and (B) MGO and diesel. Inset are the DMA-APM effective density data. The upper tails of the total (red; SMPS) distributions do not overlap the rBC distributions (black, blue) because of significant uncertainties in volume calculation for large, irregular particles, and because the latter is plotted against volume-equivalent rBC diameters, which are smaller than mobility diameters for coated or aspherical particles. SMPS measurements for diesel are not available. 809 810 811 812 813 814

Coating thickness [nm]

220 nm < $d_{rBC,vol}$ < 260 nm

s versus moderate or no coat

vith d_{rBC} 100 to 200 nm. (B)

20 nm. Negative coating thickness coating the represents an order of magnity

different size range are sh Coating thickness [nm]
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 $\begin{array}{c} 0.0 \\ 0.0 \\ \vdots \end{array}$!B! $d_{\text{rBC,vol [nm]}}$
 $d_{\text{rBC,vol [nm]}}$

4. (A) Number fractions of partic

1 by SP2 delay-time analysis, for BC-

ses of BC-containing particles with *a* as discussed in the text. Note that the the there fuels, due to practical _{16C,vol [nm]}
mber fracti
y-time anal
mtaining pa **Figure 4.** (A) Number fractions of particles with thick coatings versus moderate or no coatings, as determined by SP2 delay-time analysis, for BC-containing particles with d_{rBC} 100 to 200 nm. (B) SP2 coating thicknesses of BC-containing particles with d_{rBC} = 240 \pm 20 nm. Negative coating thicknesses are due to noise, as discussed in the text. Note that the histogram for HFO represents an order of magnitude more data than the other fuels, due to practical constraints. Note also that different size range are shown in (A) and (B) due to the fundamental difference of the two analyses (see text). The upper axis in (B) shows the predicted core-shell-Mie absorption enhancement for these coating thicknesses at the 780 nm wavelength of the CAPS PMssa. 815 816 817 818 819 820 821 822

Figure 5. MAC and SSA for the three fuels as a function of engine load. No trend with engine power or fuel is apparent. Only three error bars are shown in (A) for clarity. The shaded area in (B) reflects a MAC_{rBC,780nm} of 7.8 \pm 1.8 m² g⁻¹ (fit \pm standard error). 823 824 825

Figure 6. Measured absorption spectra for all engine test points, normalized to 880 nm. The black lines and shading indicate power laws (AAEs) and uncertainties of 2.2, 1.7 ± 0.2 , and 1.0 ± 0.1 , as labelled. All data are from the AE33 except the orange squares, which are MWAA measurements of filter samples (using a normalization factor of 0.96 to adjust from 850 to 880 nm using AAE=1.7; simultaneous AE33 measurements are highlighted with red squares). 826 827 828 829 830

Figure 7. (A) AAEs between 370 and 880 nm, (B) OM emission factors, and (C) brC MAC (referencing total OM) plotted against engine load for the 3 fuels. In (C), open symbols indicate the use of the median rather than the CAPS-PMssa-measured AE33 C value. Fewer MAC_{OM} data points than b_{abn} data points are reported due to some missing *C*OM measurements. Note that the increase of the MAC at 11 % engine load does not correspond to an increase of the OM emission factor. Note also the log scales. 831 832 833 834 835

Figure 8. Retrieved OM imaginary refractive indices plotted against MAC at 370 nm. The two high-836

valued outliers (square symbols) represent 11 %-load conditions (see Fig. [7\)](#page-27-1). The observed good correlation 837

- indicates that variability in *k*OM was driven by variability in the OM properties (particularly engine condi-838
- tions) rather than particle size or other Mie-model parameters. 839

Figure 9. Retrieved imaginary refractive indices as a function of wavelength, normalized to $k_{OM,370 \text{ nm}}$. Data where $k_{OM,\lambda}$ < 0.005 were omitted due to unreliable data (see error bars in Fig. [8\)](#page-27-0). Outlier type I refers to the high outlier at 660 nm; outlier type II refers to the two high- k_{OM} cases shown in Fig. [8.](#page-27-0) The labels λ^x refer to power laws with exponent *x*. A power law with *x* between −2 and −4 adequately describes the data. 840 841 842 843

Figure 10. Depiction of the importance of brown carbon absorption for a typical case at 50 % load. Here, brC is defined by the BC+brC approach as excess absorption after subtracting BC absorption via 844 845

AAE_{BC,880 nm}= 1. (top) Percent change in direct radiative forcing (DRF) due to brown carbon, relative to 846

- a reference case of non-absorbing organics. (middle) Contribution of organics and black carbon to overall 847
- absorption. (bottom) MAC_{OM} and k_{OM} used for this analysis. 848

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F1.

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