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4	Brown and black carbon emitted by a marine engine operated
5	on heavy fuel oil and distillate fuels: optical properties, size
6	distributions, emission factors
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25	Key Points:
26	• 50 % and 20 % of light absorption at 370 nm and 590 nm, respectively, was due to
27	"brown carbon" in the exhaust of an engine operated on heavy fuel oil (HFO).

•	• HFO contains very large rBC particles (mode at 640 nm rBC-volume-equivalent
	diameter).

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    HFO-brown-carbon λ-dependent imaginary refractive indices and mass absorption
    cross-sections are much higher than previously estimated.
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28 29

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32 Abstract

We characterized the chemical composition and optical properties of particulate mat-33 ter (PM) emitted by a marine diesel engine operated on heavy fuel oil (HFO), marine gas 34 oil (MGO) and diesel fuel (DF). For all three fuels, ~ 80% of submicron PM was organic 35 (and sulfate, for HFO at higher engine loads). Emission factors varied only slightly with 36 engine load. Refractory black carbon (rBC) particles were not thickly coated for any fuel; 37 rBC was therefore externally mixed from organic and sulfate PM. For MGO and DF PM, 38 rBC particles were lognormally distributed in size (mode at $d_{\rm rBC} \approx 120$ nm). For HFO, 39 much larger rBC particles were present. Combining the rBC mass concentrations with in-40 situ absorption measurements yielded an rBC mass absorption coefficient MAC_{BC.780 nm} 41 of 7.8 \pm 1.8 m² g⁻¹ at 780 nm for all three fuels. Using positive deviations of the absorp-42 tion Ångström exponent (AAE) from unity to define brown carbon (brC), we found that 43 brC absorption was negligible for MGO or DF PM (AAE(370,880 nm) $\approx 1.0 \pm 0.1$) but 44 typically 50 % of total 370 nm absorption for HFO PM. Even at 590 nm, ~ 20 % of the 45 total absorption was due to brC. Using absorption at 880 nm as a reference for BC absorp-46 tion and normalizing to organic PM mass, we obtained a $MAC_{OM,370 nm}$ of $0.4 m^2 g^{-1}$ at 47 typical operating conditions. Furthermore, we calculated an imaginary refractive index of 48 $(0.045 \pm 0.025)(\lambda/370 \text{ nm})^{-3}$ for HFO PM at 370 nm> $\lambda > 660 \text{ nm}$, more than twofold 49 greater than previous recommendations. Climate models should account for this substan-50 tial brC absorption in HFO PM. 51

⁵² Plain language summary (JGR feature)

We characterized the fundamental properties of marine-engine exhaust that are relevant to its aerosol-radiation interactions in climate models. In particular, we focussed on "brown carbon" light absorption (i.e., absorption in excess of that expected for the black carbon in canonical soot). We found that brown carbon can increase the 370 nm direct radiative forcing of heavy-fuel-oil marine exhaust by 18 % over snow.

58 **1** Introduction

Ship engines are the major source of combustion aerosols in the marine environ-59 ment, where almost no other anthropogenic pollution sources exist. Understanding the 60 climate-relevant optical properties of this particulate matter (PM) is therefore essential to 61 the accurate determination of anthropogenic radiative effects on the marine atmosphere 62 [Endresen, 2003; Lauer et al., 2007; Unger et al., 2010; Lindstad et al., 2015; Marelle 63 et al., 2016]. In addition, a large fraction of ship PM is emitted near shore [Eyring et al., 64 2010], which has a strong potential to negatively impact human health [Corbett et al., 65 2007; Oeder et al., 2015] as well as visibility. 66

- ⁶⁷ Ship-engine PM contains significant amounts of organic PM (OM) as well as black
- carbon (BC) [Lack et al., 2009; Buffaloe et al., 2014; Mueller et al., 2015; Price et al.,
- ⁶⁹ 2016]. When heavy fuel oil (HFO) is used, sulfates also contribute significantly to PM

mass [Lack et al., 2009; Popovicheva et al., 2009; Mueller et al., 2015]. Typically, BC con-

tributes a relatively small fraction to the mass of HFO PM. The overall single-scattering

albedo (SSA) of HFO PM aerosols is therefore relatively high, such that ship-engine PM

may result in an overall cooling over the (low-albedo) ocean surface [*Lauer et al.*, 2007;

Unger et al., 2010], but a localized warming over high-albedo surfaces or on snow and ice

⁷⁵ [*Lack and Corbett*, 2012]. The accurate quantification of these radiative effects requires an

accurate measurement of the light-absorption properties of both BC and OM.

⁷⁷ Light-absorbing OM has previously been inferred to contribute significantly to total ⁷⁸ HFO PM absorption at shorter visible wavelengths λ , based on the observation of 2- λ ab-⁷⁹ sorption Ångström exponents (AAE(370, 880 nm); Eq. 1) much greater than unity [*Mueller* ⁸⁰ *et al.*, 2015]. Accounting for this light-absorbing OM in radiative models of the Earth's at-⁸¹ mosphere requires knowledge of intensive optical properties, such as imaginary refractive ⁸² index (k_{OM}) and mass absorption cross-section (MAC_{OM}), which have not previously been ⁸³ reported.

A major goal of the present work was to quantify k_{OM} and MAC_{OM} for use in cli-84 mate models. This quantification was achieved by applying the popular [Fialho et al., 85 2005; Sandradewi et al., 2008; Favez et al., 2009; Grenfell et al., 2011; Bahadur et al., 86 2012; Cazorla et al., 2013; Lack and Langridge, 2013; Yuan et al., 2016; Pokhrel et al., 87 2017; Zotter et al., 2017] two-component "BC+brC" model: (i) light absorption in the 88 near infrared ($\lambda \ge 780$ nm) is attributed to BC only, (ii) BC absorption is assumed to 89 depend inversely on λ (AAE_{BC} = 1), and (iii) when measured visible-light absorption ex-90 ceeds predicted absorption by BC, the excess is attributed to light-absorbing OM (so-91 called brown carbon or brC). In this context, we define BC as refractory, light-absorbing 92 carbon with an AAE of unity between $370 \text{ nm} < \lambda < 950 \text{ nm}$ [Petzold et al., 2013]. We 93 consider BC as the sole light-absorbing species at $\lambda \ge 780$ nm and, when the aerosol AAE 94 exceeds unity, attribute the excess absorption to brC. While brC is therefore defined purely 95 from the optical properties of the aerosol, we also calculate intensive brC properties by 96 normalizing brC absorption to total organic PM mass. 97

For HFO PM, the two-component BC+brC absorption approach may not fully de-98 scribe the aerosol. First, HFO exhaust may contain asphaltenes, large polyaromatic molecules 99 that may contribute to PM light absorption in the near infrared, besides BC [Mullins, 100 2010]. Second, light-absorbing char particles may also be present in HFO PM [Lyyrä-101 nen et al., 1999; Chen et al., 2005; Popovicheva et al., 2009]. These char particles are 102 formed when fuel droplets graphitize rather than vapourize at the high flame temperatures 103 [Linak et al., 2000] and are therefore a subcategory of BC, with typical physical diame-104 ters of $\sim 1 \,\mu$ m. Their large physical diameter places char particles outside of the Rayleigh 105 regime, potentially invalidating the $AAE_{BC} = 1$ assumption of the BC+brC approach. Un-106 fortunately, our present measurements do not provide sufficient information to address 107 these char particles explicitly. To the extent that $AAE_{char} \neq 1$, our approach may over-108 estimate or underestimate the absorption properties of brC. However, even if this was the 109 case, our two-species (BC+brC) approach would still provide the optical parameters re-110

quired for climate models to correctly represent the wavelength-dependent light absorption of HFO PM which we have observed; the inaccuracy would be in the relative attribution of absorption to the different light-absorbing species. Moreover, as demonstrated by *Saleh et al.* [2016], climate models which employ Mie theory in radiative transfer calculations produce more accurate estimates of radiative forcing when Mie-theory-retrieved parameters are used, compared to more-complex retrievals.

In the following, we determine the optical properties of brC from a ship engine 117 operating on three different fuels: HFO, marine gas oil (MGO), and EN 590 diesel fuel 118 (DF). The upcoming sections discuss (1) speciated PM emission factors and the PM mix-119 ing state, (2) the PM SSA and MAC_{BC} at 780 nm obtained by the extinction-minus-scattering 120 technique combined with refractory BC (rBC) mass concentrations measured by a sin-121 gle particle soot photometer (SP2) and (3) brC absorption using the BC+brC conceptual 122 model in terms of MAC_{OM} and k_{OM} by using OM concentrations measured by an aerosol 123 mass spectrometer (AMS). Note that the analysis in part (3) does not incorporate the SP2 124 measurements. 125

126 2 Methods

127

2.1 Experimental

The engine used in this study was a single-cylinder research engine installed at the 128 Institute of Piston Machines and Internal Combustion Engines at the University of Ros-129 tock in Germany. The size of the combustion chamber and layout of the engine is typi-130 cal of engines used on smaller ships as a main power supply, on large ships for ancillary 131 power, or as a backup power supply on land, e.g. in hospitals. The engine is of a four-132 stroke, single-cylinder design, with a 150 mm bore and 180 mm stroke and operates at a 133 nominal 1500 rpm with a maximum power of 80 kW. Previous publications have provided 134 further details on the engine [Oeder et al., 2015; Streibel et al., 2017] and reported de-135 tailed characterizations of its emissions [Mueller et al., 2015], however, we emphasize that 136 the engine was reconfigured between these two studies. During this study, certain engine 137 operating parameters were varied to investigate their influence on the emissions; none of 138 these parameters were identified as having influenced the properties discussed herein. The 139 measurements took place between November 10th and December 9th, 2014. 140

The sampling system was similar to that used by Streibel et al. [2017] and is shown 141 in Fig. 1. Emissions (sampled 1.5 m from the engine at $> 500 \,^{\circ}\text{C}$) were first passed through 142 300 °C heated lines to a pre-cyclone with a nominal cut-off aerodynamic diameter of 2.5 μ m, 143 then diluted by a factor of roughly 12 in a two-stage dilutor. The two-stage dilutor com-144 bined a porous-tube stage, where compressed air flowed through pores in a cylinder to 145 provide a sheath flow and minimize wall losses, with an ejector dilutor, at a flow rate of 146 $150 \,\mathrm{L\,min^{-1}}$. The exact dilution ratio was measured by online CO₂ monitors. After this 147 dilution stage, filter samples for thermal/optical reflectance analysis (resulting in NIOSH 148 5040 EC/OC data) and elemental analysis (discussed elsewhere; Corbin et al. 2017) were 149

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taken. Subsequently, two tenfold ejector dilutors (Dekati Ltd., Finland) were employed to reduce PM concentrations to atmospheric levels and ambient temperatures. The steel sampling line (10 mm i.d.) to the aethalometer was approximately 5 m long with a total flow of 4.5 L min⁻¹, while the other online instruments sampled at a total of $13 L min^{-1}$ through first ~ 7 m section of tubing and then at $3 L min^{-1}$ through a second 3 m section of tubing. This flow was then divided between similar lengths of line for the instruments.

The majority of the instruments sampled at high time resolution (1 min or faster), with the exception of the particle-filter samples. For the data presented below, the 30 minute filter samples were generally used to define periods of interest (after manual inspection) and data were averaged over these periods.

Particle size distributions were obtained by a Scanning Mobility Particle Sizer (SMPS; 160 ⁸⁵Kr bipolar charger; PSI-constructed DMA equivalent to the model 3081 of TSI Inc., 161 USA; TSI CPC 3022A) and Fast Mobility Particle Sizer (FMPS, model 3091, TSI Inc.). 162 The SMPS measurements are considered more reliable as data-inversion issues (influenc-163 ing both sizing and counting) have been reported for the FMPS [Zimmerman et al., 2015]. 164 We therefore used the SMPS data to obtain geometric standard deviations (GSD) of the 165 size distribution for the Mie model. However, due to limited coverage of these SMPS 166 data, the FMPS data were used to obtain geometric mean particle diameters (GMDs). 167 These were retrieved from the FMPS data by fitting a bimodal lognormal function to the 168 data using custom code written in Igor Pro (version 6.32, WaveMetrics, OR, USA). The 169 FPMS data were first corrected using the routine published by Zimmerman et al. [2015], 170 which uses a separate CPC 3022A as a reference for the overall number concentration. 171 The GMD obtained from the corrected FMPS data is considered more reliable than the 172 FMPS GSD [Zimmerman et al., 2015]. 173

For some samples, aerosol particle mass was measured for mobility-size-selected 174 particles using a home-built DMA and a CPC (Model 3022A, TSI Inc. USA) sampling 175 behind an Aerosol Particle Mass analyzer (APM Model 3601, Kanomax Japan). The DMA-176 APM-CPC data was analyzed as described in Zieger et al. [2017] to yield particle effec-177 tive density measurements for the ultrafine mode, which was dominated by organics and 178 sulfate. Considering that the organic-sulfate ultrafine particles were liquid and therefore 179 spherical, these effective densities correspond directly to particle densities. Effective den-180 sity measurements of rBC were attempted, but were not feasible as the rBC number con-181 centrations were much lower. 182

183

2.2 Single Particle Soot Photometer (SP2)

A Single Particle Soot Photometer (SP2; Droplet Measurement Technologies, CO, USA) was employed to measure black carbon concentrations by laser-induced incandescence [*Stephens et al.*, 2003; *Schwarz et al.*, 2006]. Data were analyzed using the PSI SP2 Toolkit, version 4.112. The SP2 brings BC-containing particles to incandescence during their passage through a continuous-wave, intracavity, 1064 nm laser. The instrument detects rBC cores with mass (or volume-equivalent size, d_{rBC} , considering a void-free mate-

rial density of 1800 kg m^3) from ~ 0.7 fg (~ 80 nm) to ~ 200 fg (~ 600 nm). For parti-190 cles in this size range, the total rBC mass reported by SP2 has been validated as accurate 191 by multiple independent studies Slowik et al. [2007]; Kondo et al. [2011]; Laborde et al. 192 [2012a]. The total mass concentration reported below is corrected for the mass fraction 193 outside of this range (Section 3.1.1). We use the term "rBC" as recommended by Petzold 194 et al. [2013] whenever reporting SP2 data in a quantitative manner, e.g. C_{rBC} for BC mass 195 concentrations and d_{rBC} for mass-equivalent diameters. We use the term rBC core to clar-196 ify that any non-refractory, internally-mixed material is not measured by incandescence as 197 it vaporizes well below the ~ 4000 K sublimation point of BC. 198

The SP2 also collects light-scattering signals, acting as an optical sizer for BC-free 199 particles ($d_{\text{minimum}} \approx 160 \text{ nm}$). For BC-containing particles, the scattering signal requires 200 careful data processing because laser heating causes coatings to evaporate during measure-201 ment. This evaporation is observable in the time-resolved scattering signals. For uncoated 202 or moderately-coated particles, peak scattering signals occur when particles are relatively 203 closer to the centre of the Gaussian profile of the SP2 laser (since incident light inten-204 sity is highest closer to the centre) and simultaneously with particle incandescence. For 205 thickly coated particles, this peak is observed substantially earlier and prior to particle in-206 candescence, corresponding to the substantial reduction in particle volume when coating 207 material evaporates. This binary mixing-state classification is commonly referred to as 208 "delay-time analysis" [Moteki and Kondo, 2007]. On this basis, particles can be classified 209 as either 'thickly-coated' (significant difference between peak scattering and incandescence 210 signals) or 'moderately- or uncoated' (no significant difference). The lower detection limit 211 for this classification is higher than that for rBC quantification, since the SP2 scattering 212 measurement is less sensitive than the incandescence measurement. The upper limit for 213 this classification is imposed by saturation of the scattering signal detector, which occurs 214 for large overall particle sizes. We present data only for particles within these limits. Note 215 that a particle consisting of an rBC particle coagulated with other material may appear 216 to be 'thickly-coated' in the SP2, resulting in an overestimated fraction of thickly-coated 217 particles. Note also that a precise distinction between 'moderate or no coating' and 'thick 218 coating', for example in terms of BC volume fraction, cannot be given without knowledge 219 of the thermochemical properties of the coating [Sedlacek et al., 2015]. Based on our pre-220 vious comparisons of coating-thickness to delay-time data for atmospheric BC particles 221 (unpublished work), we estimate that 'moderate or no coating' describe particles contain-222 ing 70-100 % rBC volume. 223

A more complex data analysis approach allows the coating thickness of BC-containing 224 particles to be retrieved quantitatively, as follows. By retrieving the scattering signal at 225 3% of the maximum laser intensity, at which point no coating material has yet evap-226 orated, the scattering cross-section of the total, possibly-coated, particle can be deter-227 mined [Gao et al., 2007; Taylor et al., 2015]. Conversely, by retrieving the scattering sig-228 nal after coatings have evaporated but before the onset of incandescence, the scattering 229 cross-section of uncoated rBC can be determined [Laborde et al., 2012b]. The latter value 230 can be used to calculate the rBC-mass-equivalent volume of the rBC core, given a com-231

-6-

plex refractive index of the core $m_{core} = (n, k)$. The precise value of m_{core} is not well con-232 strained and may vary between BC materials, but the empirical relationship $k \approx (n-1)$ has 233 been shown by Bond and Bergstrom [2006], as discussed further by Moteki et al. [2010]. 234 Employing this constraint, we determined m_{core} as 1.9+0.8i by fitting the rBC-core volume 235 determined by scattering to that determined by incandescence. Our approach of adjusting 236 $m_{\rm core}$ only serves as an internal calibration of the coating-thickness retrieval, and must not 237 be misinterpreted as a measurement of the BC refractive index, which would require addi-238 tional independent measurements. Our value of $m_{\rm core}$ is smaller than that typically used in 239 SP2 analysis (2.26 + 1.26*i*), but is consistent with that used by *Laborde et al.* [2012a] for 240 propane-flame soot. Using the typical m_{core} value of 2.26 + 1.26i corresponds to smaller 241 rather than larger coating thicknesses in the analysis presented below. We note that it is 242 typical to observe some negative coating thicknesses in this analysis due to random noise 243 on the single-particle level. Investigations in our laboratory have found that this noise is 244 due mainly to the variability of particle velocity through the laser beam and uncertainty in 245 determining the scattering signal at 3 % of the maximum laser intensity. The retrieval of 246 the scattering signal at 3 % of the maximum laser intensity means that the lower detection 247 limit for coating-thickness analysis is substantially higher than for delay-time analysis. 248

The SP2 was operated and calibrated as described in [*Laborde et al.*, 2012a]. Briefly, mass-selected rBC particles representative of atmospheric or diesel rBC (Alfa Aeser Inc., FS, Lot #FS12S011) were used to calibrate the peak incandescence signal and polystyrene latex (PSL) sphere standards of diameter 269 nm were used to calibrate the partial scattering cross section measurements.

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267

2.3 Aerosol Mass Spectrometer (AMS)

Particles are introduced into the High-ResolutionAerosol Mass Spectrometer (AMS) 255 by an aerodynamic lens [Liu et al., 2007] which has its maximum transmission efficiency 256 between about 100 nm to 600 nm in free-molecular aerodynamic diameter ($d_{\rm fm,a}$). Note 257 that this aerodynamic diameter is weighted by particle density [Kulkarni et al., 2011], 258 so that $d_{\text{fm},a} = 100 \text{ nm}$ corresponds to a volume-equivalent diameter of 69 nm for the 259 HFO PM density reported below. Non-refractory material in these particles is then vapor-260 ized by an 873 K porous-tungsten cone before electron ionization. The resulting ions are 261 analyzed in a high-resolution time-of-flight mass spectrometer [DeCarlo et al., 2006]. Note 262 that we have used the term "sulfate" below rather than "sulfuric acid" (the likely form of 263 sulfate, as discussed below) because sulfate and nitrate are the chemical components mea-264 sured by AMS after electron ionization. For a more detailed discussion of the AMS re-265 sults of this study, see Corbin et al. [2017]. 266

2.4 Optical measurements

The aerosol absorption coefficient was measured by the extinction-minus-scattering technique, using a Cavity Attenuation Phase Shift PM SSA monitor [CAPS PMssa, *Onasch et al.*, 2015] operating at 780 nm. This instrument provides a calibration-free measure of

extinction (manufacturer-estimated accuracy: 5%). The extinction measurement occurs in 271 an optical cell surrounded by an integrating-sphere nephelometer. This nephelometer was 272 calibrated using the extinction measurement, as described in Onasch et al. [2015], using 273 269 nm diameter polystyrene latex spheres and size-selected NH₄NO₃. We also performed 274 a calibration at the Paul Scherrer Institute using pure N2 and CO2. Overall, eight calibra-275 tions over 13 months using all of these approaches showed a relative standard deviation of 276 only 10%, indicating very stable instrument performance. Data were manually inspected 277 to ensure that baseline values of extinction and scattering (obtained periodically on filtered 278 samples) were stable. Note that during engine measurements, the CAPS PMssa was oc-279 casionally diluted by a factor of 5.1 owing to flow limitations in some experiments (not 280 presented here), which was tested and found to have no effect on the resulting data. 281

Some data were also acquired using CAPS PMssa instruments operating at 450 nm 282 and 630 nm. These data are not presented here due to issues with data quality, possibly 283 due to the difficulty of obtaining a valid baseline measurement when variable amounts of 284 absorbing gases with strong surface interactions, such as NO₂, are present. 285

An aethalometer (model AE33, Aerosol d.o.o., Slovenia) was also deployed dur-286 ing these experiments. The AE33 measures the light attenuation through a PM deposit 287 on a filter at seven wavelengths λ . The AE33 converts filter-deposit attenuation coeffi-288 cients to aerosol light absorption coefficients by applying a conversion factor (known as 289 the "C value") as detailed in Drinovec et al. [2015] and discussed further in Section S3, 290 where the calibration factor is evaluated for our samples at 780 nm. To avoid issues due 291 to pressure changes in the sampling lines due to changes in engine conditions, our AE33 292 data was reanalyzed with a constant compensation parameter [defined in Drinovec et al., 293 2015]. 294

Three filter samples were also measured with the Multi-Wavelength Absorbance An-295 alyzer [MWAA; Massabò et al., 2013, 2015] with the goal of evaluating the performance 296 of the CAPS PMssa. In particular, the wavelength dependence of the aethalometer calibra-297 tion could be confirmed as negligible using the MWAA data. This is evident in Fig. 6, as 298 discussed further in the supplement. 299

2.5 Optical calculations 300

301

$$AAE(\lambda_1, \lambda_2) = -\frac{\ln \left(b_{ATN, \lambda_1} / b_{ATN, \lambda_2} \right)}{\ln \left(\lambda_1 / \lambda_2 \right)}$$
(1)

where the subscripts 1 and 2 indicate two different measurement wavelengths, λ is a 302 measurement wavelength, and $b_{\text{ATN},\lambda}$ is an AE33-measured attenuation coefficient. 303

Using the AE33 data, the AAE was calculated as the two-wavelength AAE:

- Using the CAPS PMssa data, $b_{abn,\lambda}$ was calculated at $\lambda = 780 \text{ nm}$ by subtracting 304
- the measured scattering from the measured extinction coefficient. By combining the CAPS 305
- PMssa $b_{abn,780}$ with the measured AAEs, we also calculated $b_{abn,\lambda}$ at $\lambda = (370, 470, 520, 590, 660)$ nm. 306

³⁰⁷ This calculation is numerically equivalent to calibrating the aethalometer using the CAPS

²⁰⁸ PMssa (full details are given in the supplement). For some samples, CAPS PMssa data

were not available for this calibration, and the median *C*-value from the other calibrations

310 was applied.

311

The MAC of a species X at a wavelength λ , MAC_{X, λ}, was calculated as

$$MAC_{X,\lambda} = b_{abn,X,\lambda}/C_X,$$
(2)

where $b_{abn,X,\lambda}$ is the aerosol absorption coefficient corresponding to absorption

by species X. In general, brC is considered to absorb negligibly at longer wavelengths X.

[Laskin et al., 2015], so that we attribute all absorption at $\lambda \ge 780 \text{ nm}$ to BC ($b_{abn,BC,\lambda} \approx$

 $b_{abn,\lambda}$ and the equation becomes

$$MAC_{rBC,780} = b_{abn,BC,780}/C_{rBC} \approx b_{abn,780}/C_{rBC},$$
 (3)

for BC, with $C_{\rm rBC}$ measured by the SP2, corrected for the mass fraction outside of the SP2 sensitivity range (Section 3.1.1).

Light absorption was attributed to brC when the measured absorption exceeded that predicted by extrapolating measurements at longer wavelengths ($\lambda \ge 780$ nm) to shorter ones ($\lambda \le 660$ nm) using an AAE= 1. That is, brC absorption was defined by

$$b_{\rm abn,brC,\lambda} = b_{\rm abn,\lambda} - b_{\rm abn,BC,880} \left(\frac{\lambda}{880\,\rm nm}\right)^{-\rm AAE,BC},\tag{4}$$

where λ and 880 nm represent two measurement wavelengths in nm and the sub-321 tracted quantity is the AAE-based estimate of BC absorption at λ , using AAE_{BC}= 1 except 322 during sensitivity tests (see below). Eq. 4 results in a positive value of $b_{abn,brC,\lambda}$ when 323 AAE_{PM} > 1 and zero $b_{abn,brC,\lambda}$ when AAE_{PM} = 1. In no case was AAE_{PM} < 1. While 324 both $b_{abn,\lambda}$ and the AAE may be affected by substantial internal mixing [Lack and Cappa, 325 2010], we expect this effect to be negligible in our data because substantial internal mix-326 ing of brC and BC was ruled out by our measurements (Section 3.1.2). Eq. 4 was em-327 ployed to calculate $b_{abn,brC,\lambda}$ at $\lambda = \{370, 470, 520, 590, 660\}$ nm, corresponding to each 328 wavelength shorter than the reference (CAPS PMssa) $\lambda = 780$ nm. 329

As shown in Eq. 4, we generally used $\lambda_2 = 880$ nm as the reference wavelength. A sensitivity test using $\lambda_2 = 950$ nm changed the calculated brC absorption by HFO by only < 2% on average. (HFO was the only fuel with AAE_{PM} substantially higher than unity.) Uncertainties in the use of Eq. 4 are discussed in Section 2.7. The MAC of brC was calculated using Eq. 2 with $b_{abn,brC,\lambda}$ and C_{OM} substituted for $b_{abn,X,\lambda}$ and C_X , respectively:

$$MAC_{OM,\lambda} = b_{abn,brC,\lambda}/C_{OM},$$
(5)

where C_{OM} is the OM mass concentration measured by the AMS.

2.6 Mie model

A Mie model was constructed and fitted to the brC absorption and OM mass concentration measurements to obtain imaginary refractive indices of OM, k_{OM} . The model structure, inputs and outputs are depicted in Fig. 2 and described in the following.

As described below, our measurements indicated that these ship-engine emissions can be appropriately modelled as an external mixture of two modes, a larger rBC mode and a smaller nucleation mode (internally mixed OM and sulfate). In estimating $b_{abn,brC,\lambda}$ from Eq. 4, we have already subtracted light absorption by BC from the total, and must only model the brC absorption when fitting the Mie model.

The model therefore consisted of a single lognormal distribution of internally-mixed 346 sulfates and OM; the reported k_{OM} values were obtained by minimizing the difference 347 between the measured absorption and the Mie-predicted absorption of this distribution 348 (each reported $k_{\rm OM}$ was fitted independently of the others). The Mie calculations required 349 several additional pieces of information, which were constrained as follows. The GMD 350 was obtained from bimodal lognormal fits to the FMPS measurements and a GSD of 1.65 351 was obtained from the SMPS data as described in Section 2.1. The respective fractions 352 of OM and sulfate in the particles were obtained from the AMS data. The particle mate-353 rial density was obtained from the DMA-APM measurements described above. For HFO 354 and MGO PM, the respectively measured material densities were $1460 \pm 50 \text{ kg m}^{-3}$ and 355 $900 \pm 40 \text{ kg m}^{-3}$. The higher density for HFO PM is due to the presence of sulfate. The 356 density measurements of MGO PM were also used to estimate the density of the DF PM, 357 which was not measured. The measured densities are in good agreement with the density 358 estimated from the mass spectra of these samples [Kuwata et al., 2012], as described in 359 the supplement. 360

The real refractive index of OM, n_{OM} was taken as n = 1.5 [Lu et al., 2015]; the 361 fitted results did not change when varying this value as $n_{\rm OM} \pm 0.1$. The corresponding 362 $n_{SO^{2-}}$ was taken as 1.35 (sulfuric acid). The overall refractive index of the particles was 363 obtained by combining $n_{\rm OM}$ and $n_{{\rm SO}_4^{2-}}$ according to their AMS-measured volume ratios, 364 following the volume-weighted linear mixing rule validated by Abo Riziq et al. [2007]. 365 Sensitivity tests indicated that the most sensitive input parameter in this model was the 366 GMD: a large change in GMD of 20 nm corresponded to a change of only ± 0.001 in 367 k_{OM} (k_{OM} was greater than 0.01 for all HFO samples, as shown in Fig. 8 below). Varying 368 other parameters by large amounts (relative to their anticipated uncertainties) generally led 369

to < 10% change in k_{OM} . We note that the PM density used here was well-constrained by the APM measurements.

³⁷² Overall uncertainties in k_{OM} were calculated by propagating the numerical uncertainties given below in a Monte Carlo calculation.

2.7 Statistical uncertainties

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³⁷⁵ Unless specified otherwise in this section, uncertainties are statistical imprecisions ³⁷⁶ and are reported as the standard error of the mean ($\sigma_{\bar{x}} = SD/\sqrt{(N)}$, SD = standard devi-³⁷⁷ ation) with the averaging periods defined by the 30 minute filter periods mentioned above. ³⁷⁸ For derived quantities, uncertainties were propagated from these $\sigma_{\bar{x}}$.

In the special cases where uncertainties other than statistical noise were known to be significant, those uncertainties were added in quadrature (i.e., possible covariance between these different uncertainties was not addressed):

• For the CAPS PMssa, uncertainties in b_{abn} were propagated from an estimated 5 % 382 accuracy in bext [as estimated by Onasch et al., 2015], 10% accuracy in scatter-383 ing calibration (based on the SD of multiple calibrations), ±1 Mm⁻¹ baselining 384 accuracy and 5 % accuracy in dilution correction. The fraction of scattered light 385 not collected by the CAPS PMssa integrating sphere (the truncation error) must be 386 corrected for if the phase function of measured particles is substantially different 387 to that of the calibration particles. This is not the case if both measured and cali-388 bration particles are small relative to the 780 nm wavelength of the CAPS Onasch 389 et al. [2015], which was generally true. Any potential complications introduced by 390 the larger BC mode in HFO PM are unlikely, considering the fact that the HFO 39 data were not outliers in the optical analyses described below. We note that the cal-392 culated absorption is not highly sensitive to truncation errors at low SSA values 393 (~0.7.) 394

• For the SP2, the relatively low number of rBC particles measured in each sampling 395 period yielded substantial statistical imprecision, which is the major source of re-396 ported uncertainty. Calibration-related uncertainties were not propagated into the 397 reported uncertainties; the precision of our calibration data was much less than the 398 An unknown source of uncertainty is caljust-mentioned statistical imprecision. 399 ibration bias related to differences between calibrant and sample. To the extent 400 that marine-engine rBC is similar to road-vehicle rBC, the latter bias is estimated 401 as < 14% [Laborde et al., 2012b]. 402

For the AMS, a measurement precision of 0.3 μg m⁻³ was estimated from measurements of filtered air and combined with the statistical imprecision of the mean. Measured concentrations may have been biased low by ~ 10 to 35 %, for smaller particles, due to the small measured particle sizes (e.g. ~ 60 nm mode diameter of volume distributions). This size-dependent bias was discussed by *Liu et al.* [2007].
Other known biases [*Jimenez et al.*, 2016] include variability in the electron ion-

409	ization efficiency of different organic molecules and the partial loss of particles
410	to bounce off of the heated vaporizer. The former effect was treated as negligible
411	and the latter estimated as negligible since these particles contained a substantial
412	amount of lubrication oil [Eichler et al., 2017], which as a liquid is unlikely to
413	bounce [Jimenez et al., 2016].
414	• An estimated inaccuracy of $\sigma_{AAE(BC)} = 0.1$ was propagated into the calculated
415	uncertainties of $b_{abn,brC,\lambda}$.

• In calculating k_{OM} , estimated imprecisions of 5 nm in GMD, 0.1 in GSD, and 4 % in density (from the SD of multiple measurements) were considered as Gaussian and propagated by Monte Carlo calculation.

419

3 Results and discussion

420

3.1 Engine emissions

The emission factors relative to engine power (in units of mg / kWh) for OM, sul-421 fates, rBC and nitrates from this engine were similar to those reported by Mueller et al. 422 [2015] for this engine. For all three fuels, OM emissions dominated the total PM mass 423 with little sensitivity to engine load. For HFO, OM contributed 58 % of the $PM_{2.5}$ mass at 424 50 % load (because sulfate contributed 25 %) and > 82 % of the PM_{2.5} mass below 50 % 425 load. Notably, when normalized to CO_2 rather than engine power, the variability in emis-426 sion factors for rBC was substantially reduced. Section S1 provides more detail on these 427 emission factors. 428

429

3.1.1 Total & BC size distributions

Particle number and volume size distributions, as measured by SMPS as a function of mobility diameter d_{mob} , are shown in Fig. 3. Also shown in these figures are the number and volume distributions of rBC cores, as measured by SP2 as a function of d_{rBC} . These size distributions are representative averages of selected measurement periods, where both SMPS and SP2 data were available.

The size distributions of the diesel fuel (DF) and marine gas oil (MGO) PM, shown 435 in Fig. 3, were comparable to that of typical diesel engines [Burtscher, 2005], in that two 436 separate modes are identifiable, a volatile ultrafine mode (at about 60 nm d_{mob} in the total-437 volume distribution) and a soot mode (lognormal at about $120 \text{ nm } d_{\text{rBC}}$ in the rBC-volume 438 distribution) were observed. AMS measurements of the fresh and thermodenuded aerosol 439 indicated that the ultrafine mode for these distillate fuels consisted mainly of organics and, 440 for MGO, sulfates. The rBC volume distribution indicated that a fraction of the MGO and 441 DF particles was above or below the quantification range of the SP2, corresponding to an 442 unmeasured mass fraction of < 3% (above) and < 23,% (below), respectively. The rBC 443 mass concentrations reported from these SP2 data have been corrected for this missing 444 fraction, as discussed further at the end of this subsection. 445

In considering Fig. 3 relative to previous studies [Oeder et al., 2015; Mueller et al., 446 2015; Streibel et al., 2017], it is important to realize that the SP2-measured $d_{\rm rBC}$ is a vol-447 ume equivalent diameter, which is by definition smaller than $d_{\rm mob}$ for soot particles of 448 both open and compact morphologies [Kulkarni et al., 2011]. For the two size distri-449 butions in Fig. 3 to overlap, a size-independent shape factor of ~ 2.2 is required. This 450 value is within the range of values (1.5–3) recommended by Sorensen [2011] for uncoated, 451 fractal-like BC aggregates, and corresponds to ~ 50 primary particles in a typical BC ag-452 gregate, if the monomer diameter is 30 nm [this monomer diameter is from Oeder et al., 453 2015]. Thus, our measured size distributions are consistent with the evidence presented in 454 Section 3.1.2 to demonstrate that rBC was externally mixed. 455

The size distribution of the residual fuel, HFO, also showed an ultrafine mode con-456 sisting of sulfates and organics with a mode diameter of about 90 nm. This mode dom-457 inated the overall particle number and had a mean (\pm SD) effective density of 1460 \pm 458 50 kg m⁻³. The effective density was considerably lower for MGO, at 890 \pm 30 kg m⁻³ 459 (data shown in Fig. 3). The effective densities for DF PM was not measured. Since the 460 particles consisted of organics and sulfate (see also Section 3.1.2) they were liquid [Se-461 infeld and Pandis, 2012], and these effective densities correspond directly to the material 462 density of this PM mode. We could not obtain effective density measurements of rBC-463 containing particles due to their relatively low number concentrations. 464

The rBC-core distribution showed a much broader volume distribution than is ex-465 pected for vehicular [Burtscher, 2005; Laborde et al., 2012c], biomass burning [Schwarz 466 et al., 2008; Laborde et al., 2012c; Wang et al., 2016], or aircraft [Lobo et al., 2015] emis-467 sions. Two studies of Chinese air pollution have also reported large rBC modes of 690 nm 468 [Huang et al., 2011] and 610 nm [Wu et al., 2017], which may be linked to HFO-combustion 469 pollution. In HFO exhaust from a marine diesel engine, Kasper et al. [2007] also observed 470 a large, non-volatile (at 400°C) mode, but did not measure its chemical composition. Our 471 measurements show that at least part of this mode consists of rBC. It likely comprised the 472 char particles defined in the introduction. Future studies should further explore the proper-473 ties of this large rBC mode. 474

The SP2 did not measure the tail of the hypothesized char mode (larger rBC mode). In the laboratory, *Linak et al.* [2000] found that char particles produced from HFO had a typical diameter of 1 μ m, which may correspond to the significant mass fraction of PM that has been identified in the coarse mode of marine engines [*Fridell et al.*, 2008]. Therefore, we initially hypothesized that the SP2 may not have measured all BC mass in the aerosol. However, we did not find supporting evidence for this hypothesis, as discussed in the following paragraphs.

A simple test of this hypothesis may be performed using the SP2 data directly. At the upper limit of single-particle mass, single particles are measured as saturated signals in the SP2. For the HFO size distribution presented in Fig. 3 only 5 of 350,000 particles caused saturated signals. Therefore, the majority of the PM passing through our sampling system (which included a PM_{2.5} cyclone) and reaching the SP2 was measured. This conclusion is corroborated by the excellent correlation between CAPS PMssa and SP2
 (Fig. S1 and S2), for which data were available only for selected experiments.

Another test of this hypothesis was performed using Mie calculations of the AAE(880,950) for BC and for cenospheres. We chose this wavelength pair so that absorption by canonical brC would be negligible [*Laskin et al.*, 2015]. Cenospheres are core-shell BC particles with air cores and have been observed in HFO PM [e.g., *Chen et al.*, 2005]. They may be modelled accurately with core-shell Mie theory [*Huang et al.*, 2012], which we have done in Fig. S3. The cenospheres were modelled with a ratio of inner to outer diameters of 0.8 (the abscissa of Fig. S3 is the outer diameter).

The Mie calculations (Fig. S3) show that in the Rayleigh regime (diameter d < 30 nm) the predicted $AAE_{BC} \approx 1$ does not depend on the model, as expected. Outside of the Rayleigh regime (d > 400 nm), the model becomes important; in our calculations the $AAE_{BC} \approx 0$. The calculated $AAE_{BC} \approx 0$ is much smaller than our observed AAE(880,950) ≥ 1 (also shown in the figure), suggesting that the majority of BC absorption and therefore BC mass did not reside in unmeasured char particles.

Overall, the HFO rBC size distribution could be described by a bimodal lognor-502 mal fit with geometric mean $d_{\rm rBC}$ and standard deviations (GMD and GSD) of 150 and 503 640 nm, and 1.0 and 0.56, for the two modes. These lognormal fits imply that 44 % of the 504 total mass was likely above (versus 2% below) the particle size range detectable by the 505 SP2 and were used to correct the total reported rBC mass. The corresponding SP2-mass 506 correction factor was 1.46 ± 0.15 . This uncertainty represents inter-experiment variability 507 and was propagated into the overall uncertainty. For unimodal lognormal fits to the diesel 508 and MGO data, the correction factors were 1.247 ± 0.05 and 1.228 ± 0.05 , respectively. 509 These latter correction factors account for particles which were too small to be measured 510 by the SP2. 511

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3.1.2 External mixing of rBC and OM/sulfate

Fig. 4A and 4B investigate the coating thickness on BC particles by analyzing the 513 time-resolved SP2 scattering signals in two different ways. Fig. 4A shows that the num-514 ber fraction of particles with evidence of 'thick coatings' (where the peak scattering sig-515 nal occurs before particle incandescence, indicating the loss of a large volume of coating 516 material due to heating) was negligible except for smaller DF particles. (Since no brC ab-517 sorption was observed for DF, this exception does not influence any of the conclusions 518 made in this study.) The remainder, 'moderate or no coating", describes particles where 519 peak scattering occurred simultaneously with incandescence. This delay-time analysis pro-520 vides a relatively coarse picture of the mixing state of BC; more quantitative information 521 is provided by the coating-thickness analysis. 522

Fig. 4B shows the results of the SP2 coating-thickness analysis. For this analysis, only BC particles with spherical-equivalent diameter 240 ± 20 nm were included, due to the need to optimize both counting statistics and signal-to-noise in the SP2 detectors. ⁵²⁶ Consistent with the delay-time analysis, the coating-thickness results show that the mean ⁵²⁷ coating thickness for HFO particles was negligible, whereas small amounts of coatings ⁵²⁸ may have been present for MGO and DF. Negative coating thicknesses reflect particle-to-⁵²⁹ particle errors (biases) in the scattering measurements, which average to zero. Note that ⁵³⁰ an order of magnitude more data were available for the HFO analysis than for the other ⁵³¹ fuels analysis. In all cases, the analysis indicates that the coatings were not thick enough ⁵³² to influence the optical properties of the BC [*Liu et al.*, 2017].

Further analysis of independent size-resolved composition measurements for HFO [*Corbin et al.*, 2017] support this picture. Ash (metal oxides and/or sulfates) is also expected in the HFO PM, as discussed in *Corbin et al.* [2017]. However, with the measurements presented here, we can only infer that ash particles did not make up the majority of the total particle mass for either rBC or OM/sulfate particles.

The effective density measurements for the mobility size range 60-120 nm shown 538 in Fig. 3 also support our conclusion that the smaller mode in that figure consisted of 539 externally-mixed OM (and sulfate, in the case of HFO) . The effective densities (calculated 540 as $6m_p/[\pi d_{\text{mobility}}^3]$, where m_p is the APM-measured single-particle mass and d_{mobility} is 541 the DMA-selected diameter, as further described in the supplement) are shown by the 542 inset circles in Fig. 3. For MGO and DF, these densities should correspond to the mate-543 rial density of the OM in those fuels. Indeed, the measurements are in close agreement 544 with the densities predicted from a published empirical relationship between the elemental 545 composition and the density of OM [Kuwata et al., 2012]. That parameterization showed 546 a low bias of just 5 % (standard deviation 2 %) relative to the DMA-APM-CPC measure-547 ments. 548

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3.1.3 Relationship of BC light absorption with SSA_{780 nm} and engine load

⁵⁵⁰ BC light absorption was characterized according to the aerosol light absorption co-⁵⁵¹ efficient measured at 780 nm, $b_{abn,780nm}$, which was measured by the CAPS PMssa using ⁵⁵² the extinction-minus-scattering technique. The measured extinction and scattering coef-⁵⁵³ ficients also provided the aerosol single-scattering albedo (SSA_{780 nm}; Fig. 5A). All three ⁵⁵⁴ fuels showed a range of SSA_{780 nm} (from 0.5 to 0.9). The lower SSAs correspond to con-⁵⁵⁵ ditions where the extinction-minus-scattering method results in a more accurate calculation ⁵⁵⁶ of $b_{abn,780nm}$.

The ratio of CAPS $b_{abn,780nm}$ to SP2 C_{rBC} gives the MAC_{BC,780nm} (Eq. 3), as shown 557 in Fig. 5B. The MAC_{BC,780nm} did not vary substantially between engine loads or fuel 558 types. The SSA780 nm did vary substantially, but without a clear dependence on engine 559 load, fuel type, or indeed any other parameter we evaluated, including the GMD (geomet-560 ric mean diameter of the FMPS) and the ratio BC:non-refractory-PM. These quantities 561 were explored because it was expected that the MAC_{BC,780nm} and SSA_{780nm} would vary 562 as a function of aerosol composition (for example, the mass fraction of BC), size (since 563 scattering depends strongly on particle size), or mixing state (since internal mixing of 564 BC with other material may result in absorption enhancements). Note that although the 565

accuracy of the CAPS-estimated $b_{abn,780nm}$ is expected to decrease with increasing SSA, our data indicate that this accuracy was not a limiting factor: there was no statisticallysignificant dependence of the MAC_{BC.780nm} on the SSA.

The lack of any such observed relationship indicates that the corresponding phenom-569 ena were not the main cause of variations in the $SSA_{780 \text{ nm}}$. Moreover, the fact that varia-570 tions in $SSA_{780 nm}$ did not correspond to variations in $MAC_{BC,780 nm}$ is consistent with the 571 aerosol containing a variable fraction of externally-mixed BC. We therefore hypothesize 572 that variable amounts of externally-mixed ash particles were the main cause of the vari-573 ability in the SSA_{780 nm}. Such ash particles were identified by electron microscopy in these 574 samples [Corbin et al., 2017], but more data would be needed to quantitatively evaluate 575 this hypothesis. 576

The shading in Fig. 5B shows the result of an uncertainty-weighted orthogonaldistance regression fit to a plot of $b_{abn,780nm}$ vs C_{rBC} (the shading shows the fit \pm 95% confidence interval; Fig. S1). As there was no apparent difference in MAC_{BC,780 nm} between different fuels, all data were fit simultaneously to yield MAC_{BC,780 nm} = 7.8 \pm $1.8 \text{ m}^2 \text{ g}^{-1}$. Assuming an AAE_{BC} of 1, this corresponds to a MAC_{rBC,550 nm} of 11.1 \pm 2.6, a factor 1.48 larger than the widely-accepted MAC_{BC,550 nm} = 7.5 m² g⁻¹ recommended by *Bond and Bergstrom* [2006] for uncoated BC.

An increase in MAC due to so-called lensing effects may occur if rBC is internally-584 mixed with substantial amounts of non-absorbing material [Liu et al., 2017]. However, 585 our SP2 measurements indicate that much more internal mixing than observed would be 586 necessary to explain a 48 % lensing enhancement (Fig. 4). The calculated absorption en-587 hancement for a 240 nm rBC core is shown on the upper axis of Fig. 4, using a Mie ap-588 proximation. For HFO, a negligible enhancement is predicted. For MGO and DF, a max-589 imum absorption enhancement of 10-20% is predicted. Further calculations verified that 590 these enhancements change negligibly when the soot aggregates are represented as non-591 interacting monomer spheres, instead of volume-equivalent spheres [Liu et al., 2015]. Con-592 sidering the uncertainties associated with the coating-thickness and MAC data, we con-593 sider that the discrepancy between the calculated and apparent absorption enhancements 594 is probably negligible for MGO and DF. For HFO, where our confidence is enhanced by 595 a larger number of analyzed particles, we find that lensing does not explain the why the 596 MAC_{rBC.780 nm} was higher than expected. Further work is needed to clarify why this HFO 597 MAC_{rBC,780 nm} was higher than previously-reported values. 598

Some difference in MAC_{rBC.780 nm} for HFO compared to the distillate fuels may have 599 been anticipated due to the presence of a large rBC mode in HFO PM (Section 3.1.1). 600 This was not observed. One may therefore hypothesize that all rBC particles consisted 601 of soot aggregates, since the $MAC_{rBC,780 nm}$ of a soot aggregate depends primarily on the 602 size of its monomers [Sorensen, 2001] and not on its overall size. However, as measure-603 ments on char particles are lacking in the literature, it is not clear that the MAC_{rBC,780 nm} 604 (and other optical properties) of small char particles would be sufficiently different from 605 soot to be seen in Figs. 5B and S1. Regardless of the physical interpretation, our data 606

indicate that the anomalous size distribution of HFO PM did not lead to an anomalous $MAC_{rBC,780 \text{ nm}}$. We note, however, that SP2 measurements were not available for the low-

est engine load test points, for which brC absorption appeared to be anomalously high.

3.2 Brown carbon

In this section, we discuss the wavelength-dependence and intensive properties of brown carbon (brC). The focus is on HFO PM, since brC was negligible in the other fuels.

We reiterate here that these calculations are numerically independent of the SP2 data discussed in Section 3.1.1, as shown in the schematic diagram Fig. 2. Uncertainties related to the larger mode of rBC particles measured in HFO PM by SP2 are therefore irrelevant to the following results. In particular, we used only the $b_{abn,780nm}$ measured by CAPS PMssa and the AAE measured by the AE33 to calculate k_{OM} . The assumptions are that AAE_{BC}= 1 and that only BC absorbs at $\lambda > 700$ nm.

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3.2.1 Wavelength dependence of light absorption

Fig. 6 shows AE33 absorption spectra for all engine test points. The data are normalized to 880 nm to allow different concentrations to be compared. The distillate fuels (DF and MGO) show similar trends and are consistent with an AAE of 1 ± 0.1 down to 470 nm. At 370 nm, there is some evidence of a minor contribution (< 20 % of b_{abn} at 370 nm) of brC to light absorption for the distillate fuels.

For HFO, the data are generally consistent with an AAE(370–950) of 1.7 ± 0.2 . The shaded region in Fig. 6 encompasses this range, showing that this AAE includes the majority of the data points while slightly underpredicting absorption at $\lambda < 500$ nm. We hypothesized that the scatter in Fig 6 may be related to a varying mass fraction of BC (relative to OM and sulfate) but found this this was not the case (Fig. S4).

Fig. 6 also includes MWAA measurements of light absorption of the filter samples. These spectral measurements provide additional confidence in the AE33 data, since the MWAA directly measures and corrects for light scattering from the filter sample. The MWAA data represent only three HFO filter samples due to limited sample availability, but generally corroborate the AE33-measured AAEs.

636 3.2.2 MAC_{OM,370}

Fig. 7C shows brC absorption at 370 nm, per mass of OM $(MAC_{OM,370})$ for all fuels as a function of engine load. For cases in which the median AE33 *C* value was used, rather than a specifically-calibrated *C*-value, the data are shown with open rather than filled symbols. The upper panel, Fig. 7A, illustrates the AE33-measured AAEs from which these MACs have been derived.

Fig. 7C shows that the MAC_{OM.370} increased by an order of magnitude at the low-642 est engine load for HFO. This increase did not correspond to a change in the overall OM 643 emission factor EF_{OM}, as seen by comparison to Fig. 7B, nor to a change in the sampled 644 OM concentrations (not shown). Therefore, this change must be due to a difference in the 645 chemical composition of the OM. For this reason, we hypothesized that the less-volatile 646 thermal-optical OC fractions would correlate with MAC_{OM.370}, but found that the cor-647 relation was poor (Section S5). The relative emission factors of polycyclic aromatic hy-648 drocarbons (PAHs), including oxygenated- and nitrogenated PAHs, have been reported to 649 increase at low engine loads by Sippula et al. [2014]. Such species have been connected to 650 the chemical composition of the fuel [Rüger et al., 2015; Streibel et al., 2017]. The opti-651 cal properties are thus not adequately described by the lubricating-oil model proposed by 652 Eichler et al. [2017], who only discussed engine emissions at 50% load. 653

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3.2.3 Imaginary refractive index of OM, $k_{OM,\lambda}$

Based on the measurements described in the previous sections, a Mie model was constructed and fit to the data to obtain the imaginary refractive index of OM for this study, as detailed in Section 2.5.

Figs. 8 and 9 show the results of the Mie model fits. Fig. 8 shows $k_{OM,370}$ as a function of MAC_{OM,370 nm}. The two quantities are correlated because the main variations in model inputs were in OM mass concentration and light absorption, rather than particle size or density. Two extreme outliers have a $k_{OM,370}$ close to unity, which is unusually high. We carefully inspected all input data for these measurements but found no errors. The $k_{OM,370}$ for DF and MGO were not significantly different from zero.

Fig. 9 depicts the wavelength dependence of k_{OM} by normalizing $k_{OM,\lambda}$ to $k_{OM,370}$ for all data. The grey shading and lines show the behaviour that would be expected for a k_{OM} AAE of 2, 3, and 4 (upper line, central line, lower line). The plot was filtered to only included points where $k_{OM,\lambda} > 0.005$ before normalization. This threshold is based on the Monte-Carlo-estimated uncertainties shown in Fig. 8 and resulted in the exclusion of all MGO and DF data.

Two types of outliers are noticeable in Fig. 9. The first, "outlier I", shows an anoma-670 lously slow decrease in $k_{\rm OM}$ with λ . The reason for this exception was not identified. 671 The second, "outlier II", correspond to the two high outliers in Fig. 8. The k_{OM} -versus-672 λ trend for both of these type-II outliers appears reasonable and is similar; compared 673 to the other measurements the slope is more negative than expected. In other words, in 674 addition to having higher absolute $k_{OM,370}$ these measurements also had a higher ratio 675 $k_{OM,370}$: $k_{OM,660}$, which is consistent with a true chemical difference in the type-II sam-676 ples. 677

678 4 Discussion

679

4.1 Expected differences between engines and HFO samples

In general, the intensive optical properties of marine-engine PM may vary between studies. In our present study, we have observed similar trends to those observed in the study described by *Mueller et al.* [2015], as discussed in detail in Section S5. Further measurements are required to investigate the degree to which engine-to-engine variability may influence these properties.

In addition, differences in fuel composition may influence marine-engine PM optical 685 properties. Since the molecular composition of HFO PM has been shown to be compara-686 ble with the molecular composition of HFO [Streibel et al., 2017], in a manner depending 687 on engine load [Sippula et al., 2014], we hypothesize that the molecular composition of 688 brC is also related to the molecular composition of the fuel. In support of this hypothesis, 689 we note that Sippula et al. [2014] observed larger mass fractions of PAHs in HFO OM at 690 lower engine loads, similar to our observation of increased MAC_{OM} at lower engine loads. 691 Those authors suggested that the PAHs in HFO PM represent unburned fuel. This picture 692 is consistent with the non-absorbing nature of the distillate-fuel PM samples, since neither 693 MGO nor DF contain substantial levels of PAHs. 694

If PAHs in HFO are a major source of brC in HFO PM, then different engines may produce brC in a similar manner to one another. Thus, the extrapolation of the MAC_{OM} and k_{OM} reported herein to other engines may be justified, especially when using the relationship of OC3 / OC with MAC_{OM,370} (Fig. S5) to estimate light-absorption by brC in HFO PM. However, since variability in the composition of HFO fuel [*Stout and Wang*, 2016] may lead to variability in the light-absorbing properties of the emissions, additional studies are warranted to constrain fuel-related variability in k_{OM} and MAC_{OM}.

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4.2 Summary and implications of brC / OM light absorption

⁷⁰³ In the following, we summarize the results on brC absorption and provide an illus-⁷⁰⁴ tration of the potential radiative impacts of HFO brC.

The standard operating load for this engine is 50 %, although engines frequently operate at lower loads to save fuel or in the presence of ice [*Lack and Corbett*, 2012]. At 50 % load, the mean $k_{\text{OM},370\text{nm}}$ and corresponding SD were 0.057 \pm 0.027 (n = 5). At 25 % load, the corresponding values were 0.034 \pm 0.018 (n = 6). Considering the large variabilities, we recommend the combined mean, 0.045 \pm 0.025, for modelling studies. According to Fig. 9, $k_{\text{OM},\lambda}$ may then be estimated as

$$k_{\text{OM},\lambda} = (0.045 \pm 0.025)(\lambda/370 \text{nm})^{-3}, \qquad 370 \text{ nm} < \lambda < 660 \text{ nm}$$
 (6)

At the commonly-used λ of 550 nm, Equation 6 gives $k_{\text{OM},550\text{nm}} = 0.014 \pm 55 \%$. Compared to the $k_{\text{OM},550\text{nm}}$ values summarized by the review of *Lu et al.* [2015], this is a ⁷¹³ high value. For example, those authors recommended $k_{OM,550nm} = 0.015 \pm 45\%$ for lig-⁷¹⁴ nite OM and $k_{OM,550nm} \approx 0.03$ for biomass-combustion OM with 10% BC (a comparable ⁷¹⁵ fraction to our data). For HFO, *Lu et al.* [2015] recommended $k_{OM,550nm} = 0.006 \pm 40\%$, ⁷¹⁶ based on the extrapolation of studies of DF emissions to HFO. This is a factor of 2.3 ⁷¹⁷ smaller than our measured value. Considering that we measured negligible brC for DF ⁷¹⁸ PM, this extrapolation is not valid and future studies should use Equation 6 to estimate ⁷¹⁹ k_{OM} for HFO.

The very high OM fraction (approximately 90%) in HFO emissions makes brC light 720 absorption especially important for this aerosol in spite of its k_{OM} falling within the range 721 of previously-reported values (at 25–50 % engine load). This is demonstrated in Fig. 10, 722 where the Simple Forcing Efficiency [Chen and Bond, 2010; Bond and Bergstrom, 2006] 723 has been employed. The figure illustrates the Direct Radiative Forcing (DRF) of a rep-724 resentative ship-exhaust plume over the planet's surface. A wavelength-dependent mass 725 scattering cross section of 2.3 m² g⁻¹ at 780 nm and scaling exponent of 4 (see supple-726 ment for details) was used with a backscatter fraction of 0.1. For the 50 % load case (in-727 puts shown in lower panel), the figure plots the relative absorption of OM and BC (middle 728 panel) and the estimated change in direct radiative forcing (DRF, upper panel) when in-729 cluding brC absorption. The change in DRF is relative to the base case of non-absorbing 730 OM (equal OM mass, with k_{OM} set to zero). Relative humidity effects have been ne-731 glected for simplicity. The DRF is highly sensitive to the albedo of the surface below the 732 aerosol; as such, two cases are shown: one for the case of fresh snow or cloud, one for the 733 case of the dark ocean. 734

Note that OM absorption from HFO at $\lambda = 660$ nm is ~ 20 % of the total. At this 735 wavelength, brC absorption may have been expected to have dropped to zero [Laskin et al., 736 2015]: HFO brC clearly absorbs farther into the red than expected. At $\lambda = 370$ nm, the 737 contribution of brC to the DRF increases to 18% for the over-snow case. Note that we 738 have chosen a conservative example of 50 % load, and that MAC_{OM} was much larger at 739 lower loads (Fig. 7). The ability of brC absorption to influence the DRF of HFO-exhaust 740 PM, particularly in high-albedo environments like the Arctic, should motivate modelling 741 studies to assess the overall climate effects of this organic absorption. 742

743 **5** Conclusions

We used in-situ measurements of single-particle black carbon mass (via laser-induced incandescence), in-situ aerosol light extinction and scattering coefficients, and filter-based measurements of Ångström absorption exponents, to reach several conclusions on the properties of BC and brC in emissions from an auxiliary-scale research ship engine. One residual fuel (heavy fuel oil; HFO) and two distillate fuels (marine gas oil, MGO, and diesel fuel, DF) were used.

Overall, light absorption by OM in DF and MGO exhaust was found to be negligible, whereas absorption by OM in HFO PM was very significant. This result is not sur-

752	prising considering the near-black colour of the HFO fuel (diesel is almost transparent)
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and the molecular similarity between HFO-exhaust OM and the fuel [*Sippula et al.*, 2014;

Streibel et al., 2017]. In more detail, our conclusions are as follows.

755	• Generally, BC was externally mixed from non-refractory material (sulfates and or-
756	ganics) in the primary emissions. The BC size distribution for DF and MGO was a
756	typical lognormal, whereas for HFO an apparent bimodal lognormal with very large
758	rBC particles was observed.
	 The ratio of 780 nm light absorption coefficient with rBC mass gave a MAC_{rBC.780 nm}
759	• The faile of 780 min light absorption coefficient with FBC mass gave a MAC _{FBC,780 nm} of 7.8 \pm 1.8 m ² g ⁻¹ from SP2 and in-situ absorption measurements for all three fu-
760	els.
761	
762	• The overall wavelength dependence of absorption (370 < λ < 950 nm) for this PM
763	can be generally described by an AAE of 1.0 ± 0.1 for MGO and DF and by an
764	AAE of 1.7 ± 0.2 for HFO.
765	• Considering the PM as a mixture of brC and BC with $AAE_{BC}=1$ (BC+brC model),
766	only HFO was found to contain significant amounts of brC, with $MAC_{OM,370 nm}$ =
767	$0.4 \pm 0.2 \mathrm{m^2 g^{-1}}$ at typical operating conditions of 50 % load. At low engine loads
768	(11 %), MACs up to $10 \text{ m}^2 \text{ g}^{-1}$ were observed while OM emission factors remained
769	unchanged, indicating a fundamental change in the nature of the absorbing material.
770	• Using a measurement-constrained Mie model, we retrieved wavelength-dependent
771	imaginary refractive indices $k_{OM,\lambda}$ for HFO of $0.032 - 0.066$ (interquartile range).
772	A parameterization was provided for $k_{OM,\lambda}$ at typical engine operating conditions
773	(Eq. 6). The reported k_{OM} is large compared to previous values reported for open
774	biomass burning [Lu et al., 2015]. Due to the high OM fraction in this exhaust (~
775	90 %), OM absorption (brC) may contribute up to 50 % of total absorption for HFO
776	at 370 nm, and 20 % at 660 nm.
777	• A simplified calculation showed that accounting for brC in the radiative forcing of
778	HFO PM may enhance the direct radiative forcing by 18 $\%$ at 370 nm and 50 $\%$
779	load, and much more at lower engine loads. Climate models should therefore in-
780	clude this absorption when estimating the radiative effects of ship exhaust.

- Although the emissions and light-absorbing properties of OM may vary between engines, between engine settings, and between HFO samples, our reported k_{OM} and MAC_{OM} values provide a significant improvement over previously-available data.
- 784 Author contributions

JCC performed the initial data analysis and interpretation. HC, MG, IEH, GM, and LD contributed significantly to data interpretation. MZ, JCC and MG operated and calibrated the SP2. SP and JCC operated and calibrated the CAPS. GJ operated the AE33 and provided its data, LD analyzed the AE33 raw data. AAM, SP, and JCC operated and calibrated the AMS. JCC performed the APM measurements. JO and BS planned and oversaw the experiments. ASHP and RZ initiated the study. JCC wrote the manuscript and all

⁷⁹¹ authors contributed to the writing.

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ter AE33 was developed and is being produced. Emission factor and brC properties data

are included in the online supplementary material.

Table 1. Common abbreviations and symbols used in the text. Rarely-used abbreviations have been omitted

804 for brevity.

Abbreviation	Definition
AAE	Absorption Ångström exponent
$AAE(\lambda_1, \lambda_2)$	AAE calculated from two specified λ
AE33	Model AE33 Aethalometer
BC	Black carbon
rBC	refractory BC, measured by SP2
brC	Brown carbon, operationally defined by light
	absorption in excess of that predicted for BC.
CAPS PMssa	Cavity Attenuation Phase Shift
	PM SSA monitor.
AMS	Aerosol Mass Spectrometer
APM	Aerosol Particle Mass analyzer
DF	Diesel fuel (a distillate fuel)
MAC	Mass absorption cross-section of PM
MGO	Marine Gas Oil (a distillate fuel)
MWAA	Multi-Wavelength Absorption Analysis
HFO	Heavy Fuel Oil (residual fuel)
ОМ	The non-refractory organic component of PM,
	operationally defined by vaporization at $600 ^{\circ}\text{C}$
PM	Particulate Matter
SSA	Single-scattering albedo
SP2	Single Particle Soot Photometer
$b_{\rm abn}$	Aerosol absorption coefficient
C_x	Mass concentration of <i>x</i>
n	Real part of the refractive index
k	Imaginary part of the refractive index
λ	Wavelength
$d_{\rm rBC}$	Volume-equivalent diameter of rBC core
$d_{\rm mob}$	Mobility diameter

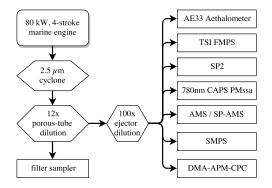
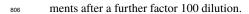


Figure 1. Experimental setup. Filter samples were measured after a factor ~ 12 dilution; other measure-



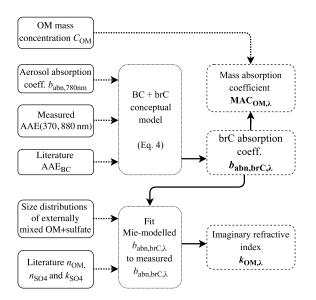


Figure 2. Schematic of the BC+brC analysis and Mie model fit performed herein. Symbols and abbreviations are defined in Table 1.

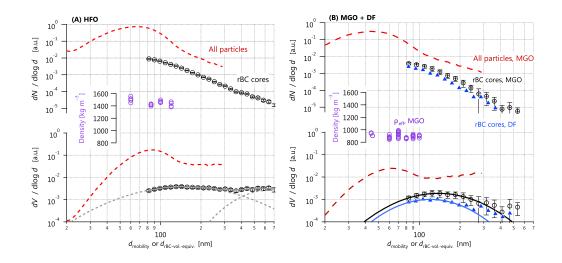
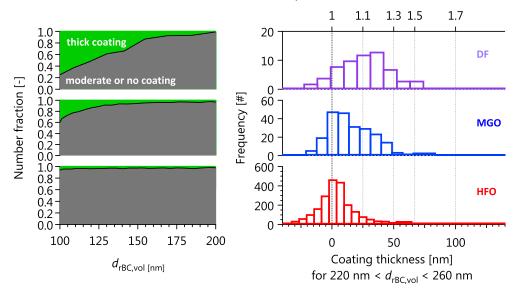


Figure 3. Size distributions of PM from the marine engine operated on (A) HFO and (B) MGO and diesel. Inset are the DMA-APM effective density data. The upper tails of the total (red; SMPS) distributions do not overlap the rBC distributions (black, blue) because of significant uncertainties in volume calculation for large, irregular particles, and because the latter is plotted against volume-equivalent rBC diameters, which are smaller than mobility diameters for coated or aspherical particles. SMPS measurements for diesel are not available.



Absorption ratio (core-shell : bare BC) 780 nm [-]

(A) Number fractions of particles with thick coatings versus moderate or no coatings, as de-Figure 4. 815 termined by SP2 delay-time analysis, for BC-containing particles with d_{rBC} 100 to 200 nm. (B) SP2 coating 816 thicknesses of BC-containing particles with $d_{\rm rBC} = 240 \pm 20$ nm. Negative coating thicknesses are due 817 to noise, as discussed in the text. Note that the histogram for HFO represents an order of magnitude more 818 data than the other fuels, due to practical constraints. Note also that different size range are shown in (A) 819 and (B) due to the fundamental difference of the two analyses (see text). The upper axis in (B) shows the 820 predicted core-shell-Mie absorption enhancement for these coating thicknesses at the 780 nm wavelength of 821 the CAPS PMssa. 822

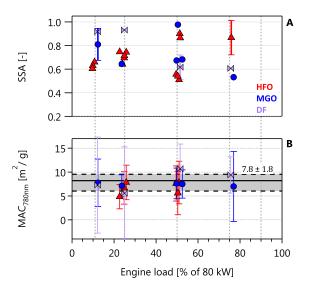


Figure 5. MAC and SSA for the three fuels as a function of engine load. No trend with engine power or fuel is apparent. Only three error bars are shown in (A) for clarity. The shaded area in (B) reflects a MAC_{rBC,780nm} of $7.8 \pm 1.8 \text{ m}^2 \text{ g}^{-1}$ (fit \pm standard error).

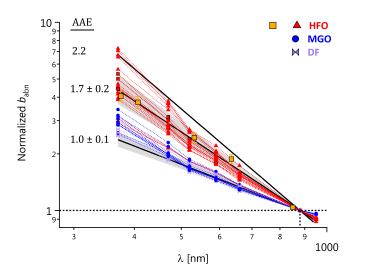


Figure 6. Measured absorption spectra for all engine test points, normalized to 880 nm. The black lines and shading indicate power laws (AAEs) and uncertainties of 2.2, 1.7 ± 0.2 , and 1.0 ± 0.1 , as labelled. All data are from the AE33 except the orange squares, which are MWAA measurements of filter samples (using a normalization factor of 0.96 to adjust from 850 to 880 nm using AAE=1.7; simultaneous AE33 measurements are highlighted with red squares).

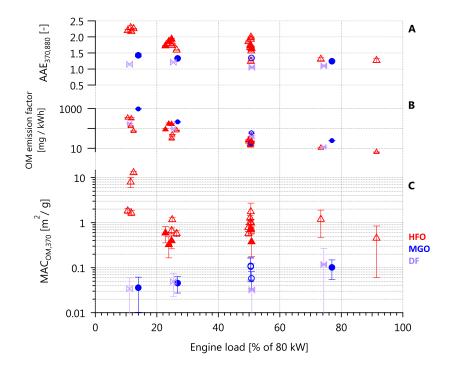


Figure 7. (A) AAEs between 370 and 880 nm, (B) OM emission factors, and (C) brC MAC (referencing total OM) plotted against engine load for the 3 fuels. In (C), open symbols indicate the use of the median rather than the CAPS-PMssa-measured AE33 C value. Fewer MAC_{OM} data points than b_{abn} data points are reported due to some missing C_{OM} measurements. Note that the increase of the MAC at 11 % engine load does not correspond to an increase of the OM emission factor. Note also the log scales.

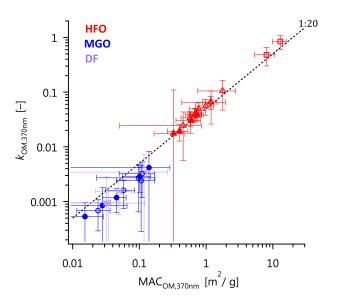


Figure 8. Retrieved OM imaginary refractive indices plotted against MAC at 370 nm. The two high-

valued outliers (square symbols) represent 11 %-load conditions (see Fig. 7). The observed good correlation

indicates that variability in k_{OM} was driven by variability in the OM properties (particularly engine condi-

tions) rather than particle size or other Mie-model parameters.

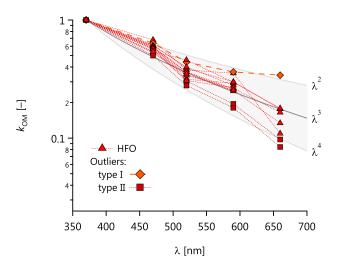


Figure 9. Retrieved imaginary refractive indices as a function of wavelength, normalized to $k_{OM,370 \text{ nm}}$. Data where $k_{OM,\lambda} < 0.005$ were omitted due to unreliable data (see error bars in Fig. 8). Outlier type I refers to the high outlier at 660 nm; outlier type II refers to the two high- k_{OM} cases shown in Fig. 8. The labels λ^{x} refer to power laws with exponent *x*. A power law with *x* between -2 and -4 adequately describes the data.

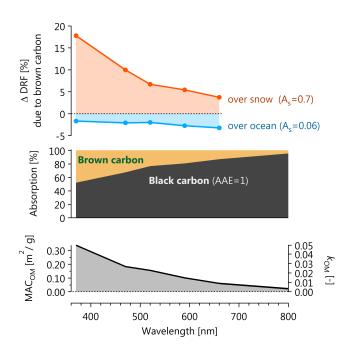


Figure 10. Depiction of the importance of brown carbon absorption for a typical case at 50 % load. Here, brC is defined by the BC+brC approach as excess absorption after subtracting BC absorption via AAE_{BC,880 nm}= 1. (top) Percent change in direct radiative forcing (DRF) due to brown carbon, relative to a reference case of non-absorbing organics. (middle) Contribution of organics and black carbon to overall absorption. (bottom) MAC_{OM} and k_{OM} used for this analysis.

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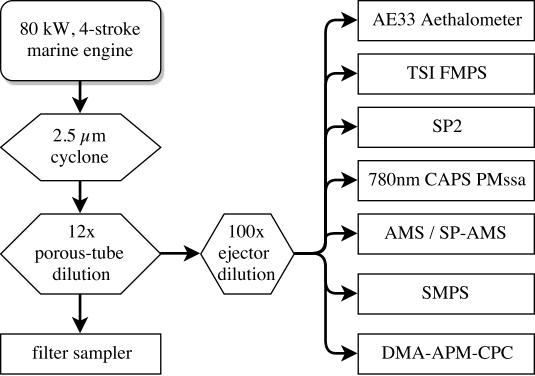
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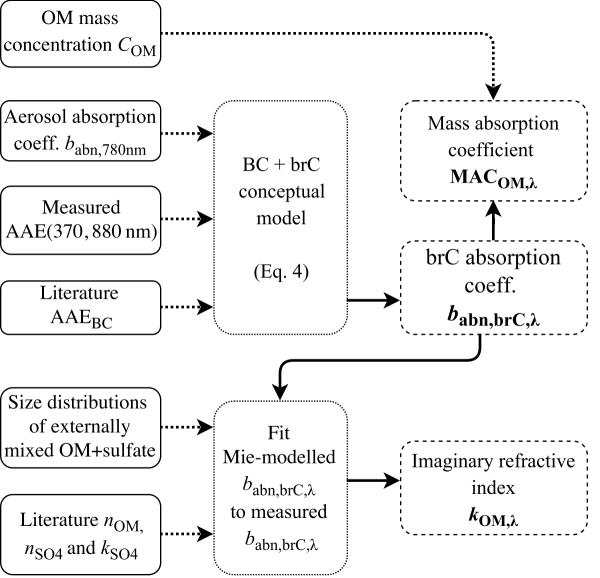
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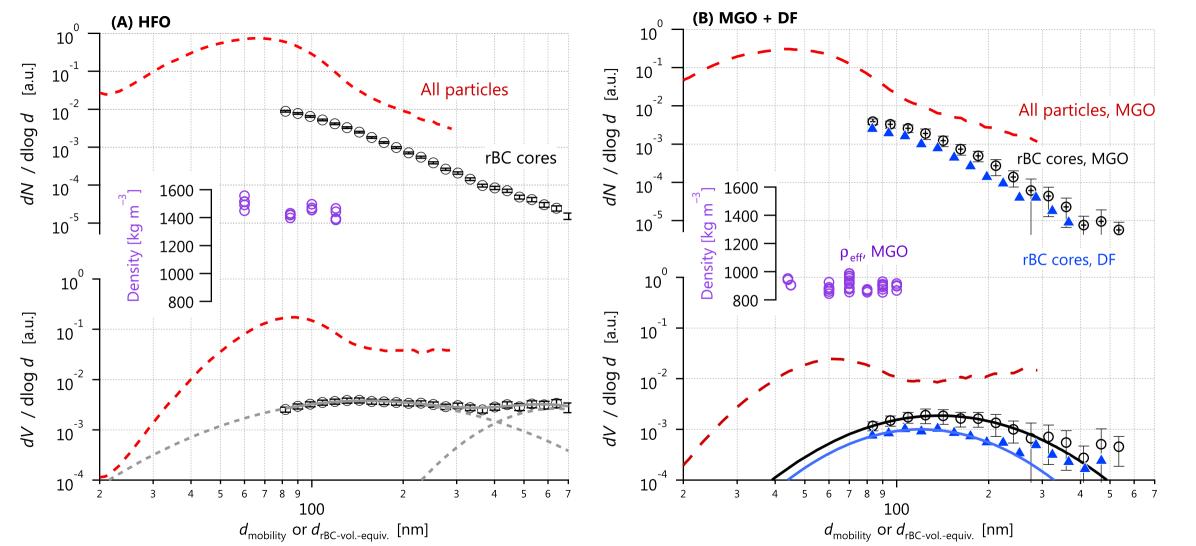
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F1.



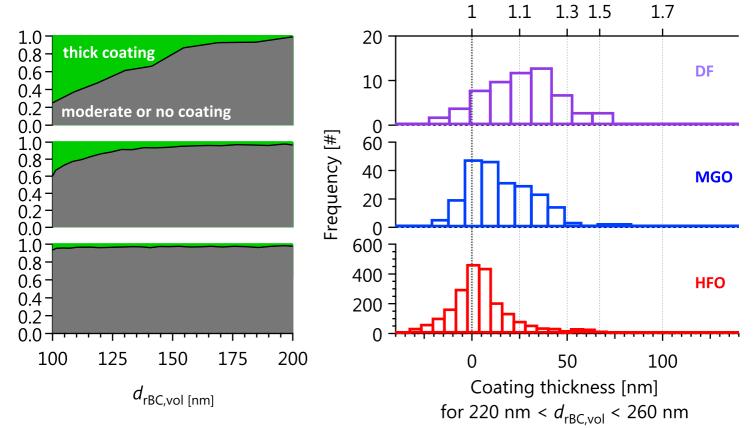


F3.



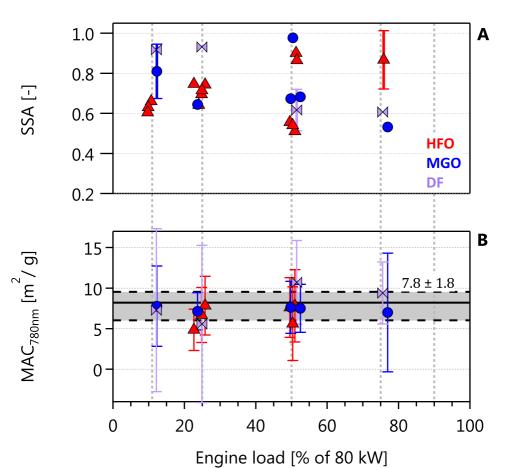
F4.

Absorption ratio (core-shell : bare BC) 780 nm [-]

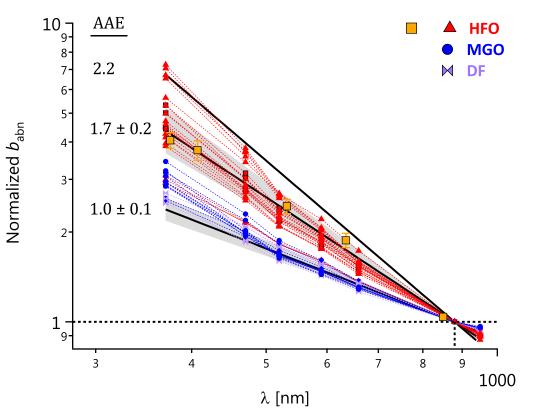


Number fraction [-]

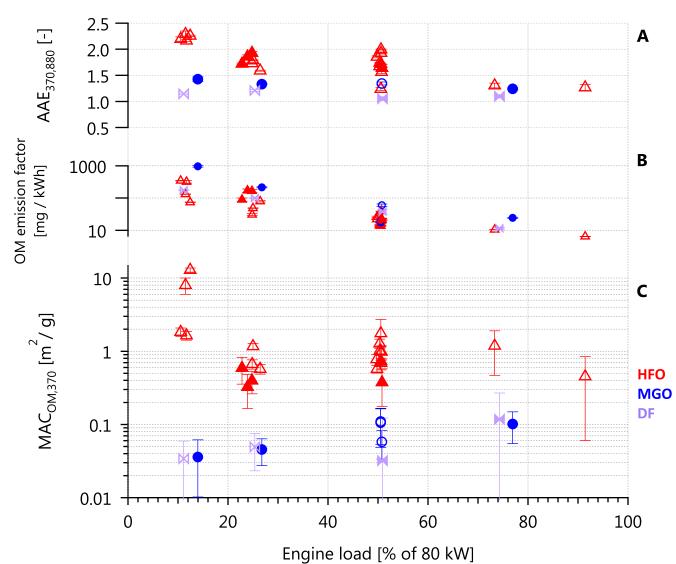
F5.



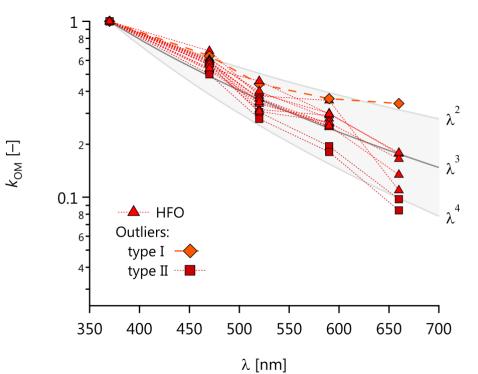
F6.



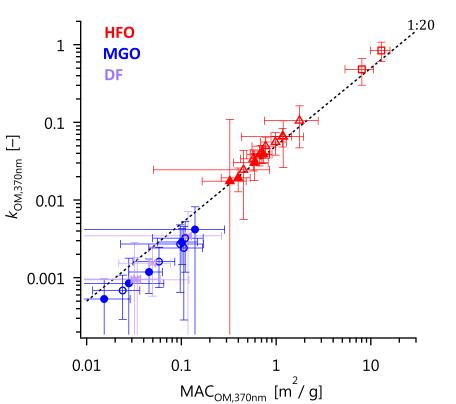
F7.



F8.



F9.



F10.

